The big bang – is it true?

Theory, models and evidence in 20th century physics







Cormac O'Raifeartaigh FRAS (WIT)

Overview

♯ The big bang model

U was once superdense and superhot Expanding and cooling over time

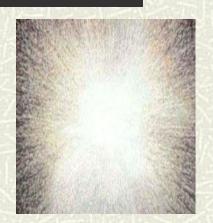
The runaway galaxies; the abundance of the elements
The distribution of the galaxies; the background radiation

The theory

Cosmology and the general theory of relativity

What does a historian do?

Review how models develop Consider roads not taken







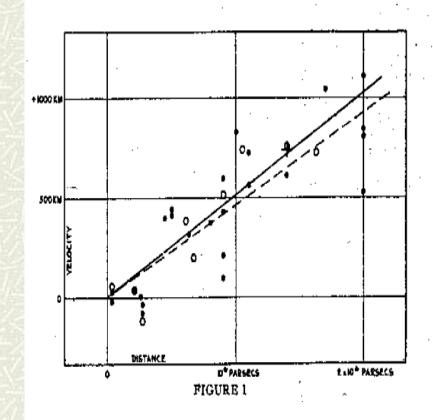


1st evidence: Hubble's law

- **♯** The recession of the galaxies
- Linear relation between redshift (velocity) and distance
- # Hubble's Law (1929)

Far-away galaxies rushing away at a speed proportional to distance

Velocity-Distance Relation among Extra-Galactic Nebulae.

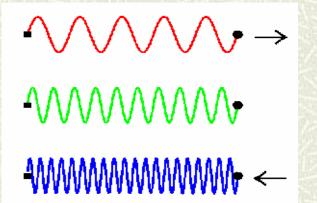


$$v = H_o d$$

Motion of galaxies: redshift



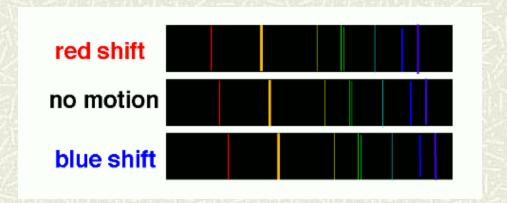
Vesto Slipher (1915)



frequency of light depends on motion of source relative to observer

Doppler Effect

measure <u>motion</u> of stars and galaxies from light emitted

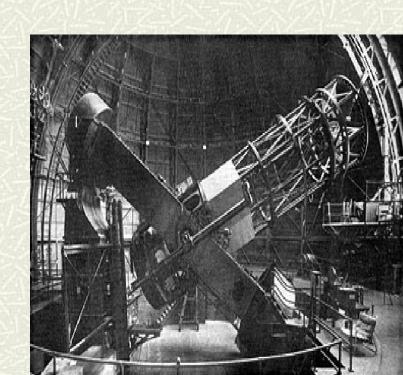


The distance of the galaxies



Edwin Hubble

- ★ Identification of Cepheid variables stars in spiral nebulae (1925)
- Use of Leavitt period-luminosity method
- ★ Nebulae far beyond Milky Way
- **♯** Distinct galaxies



Explanation for runaway galaxies?

Newton

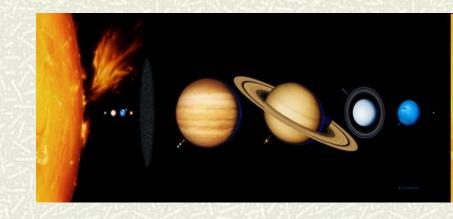
- Gravity pulls in not out
- Space is fixed

How can galaxies be receding?

What is pushing out?



Isaac Newton



Modern theory of gravity

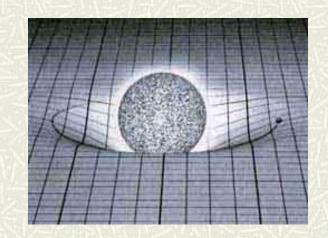
General theory of relativity (Einstein, 1915)

- space + time not fixed
- spacetime
- affected by motion
- affected by mass

$$G_{\mu\nu} = -kT_{\mu\nu}$$

gravity = curvature of space-time





Relativistic cosmology



Time-varying solutions for the cosmos Expanding or contracting universe

Evolving universe

Time-varying radius and density of matter Rejected by Einstein

♯ Georges Lemaître (1927)

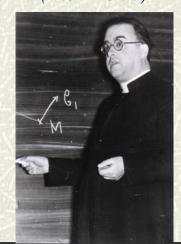
Relativistic universe of expanding radius
Agreement with emerging astronomical data
Also rejected by Einstein





Alexander Friedman (1888 -1925)

Georges Lemaître (1894-1966)



The paradigm shift (1930)

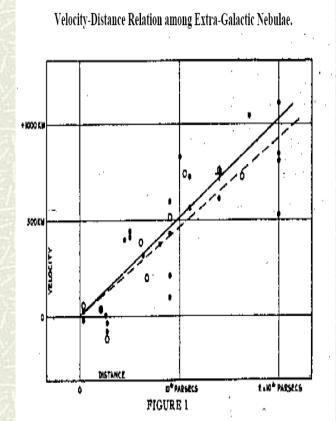
• RAS meeting (1930)

Eddington, de Sitter

<u>If</u> redshifts are velocities and effect is non-local

- Hubble's law = <u>expansion of space</u>?
 Static models don't fit data
 Dynamic relativistic models required
- Friedman-Lemaître models accepted

Time-varying radius
Density of matter decreases





Evidence now favourable

The Friedman-Einstein model

Einstein's first expanding model (1931)

Friedman model with cosmological constant set to zero First English translation



Cosmic radius $R \sim 10^8 LY$

Density of matter $\rho \sim 10^{-26}$ g/cm³

Timespan of expansion: 10¹⁰ yr

Numerical error in calculations

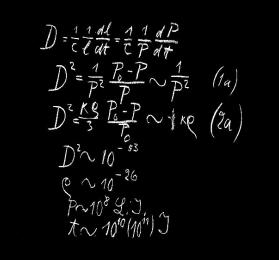
Source of error not clear

Made clear on Oxford Blackboard

Nature of error clear on blackboard





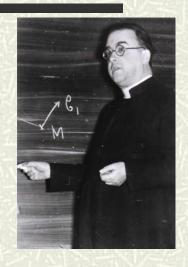


The first 'big bang' model (1931)

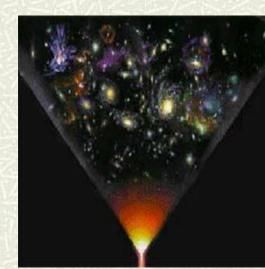
- \blacksquare Expanding *U* smaller in the past
- **■** Rewind Hubble graph to 'origin'
- # Extremely dense, extremely hot
- \blacksquare Explosive beginning at t = 0? Expanding and cooling ever since



Not accepted



Fr Georges Lemaîre



A second piece of evidence

- **#** How did the chemical elements form?
- **♯** Nuclear physics (1940s)
- **♯** Not in the stars
- **■** In Lemaître's infant universe?
- **#** *H*, *He* nuclei (1 s)
- U = 75% H, 25% He
- **♯** Agrees with observation



Georges Gamow



Big bang nucleosynthesis

A third piece of evidence?

- **#** Radiation from infant universe
- **■** Released when atoms formed (300,000 yr)
- **★** Still observable today?

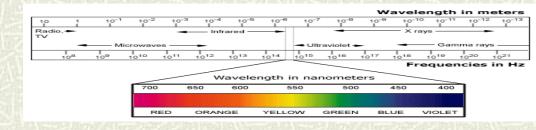
 Low temp, microwave frequency



Alpher, Gamow and Herman

No-one looked (1948); why not?





Steady-state model (1950s)



Fred Hoyle



- # Rival model
- **#** Expanding universe

BUT

- **♯** Matter continuously created
- ★ No beginning

A famous debate (1950-1965)

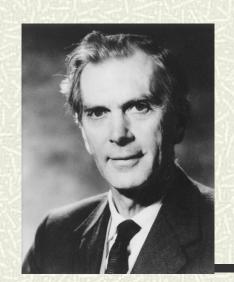
- **♯** Bitter debate between BB and SS
- **Radio-astronomy**: study most distant galaxies
- **#** Compare with local galaxies
- **■** Distribution the same at all times? (SS)
- **■** Or different? (BB)

Answer: different



End of steady-state model



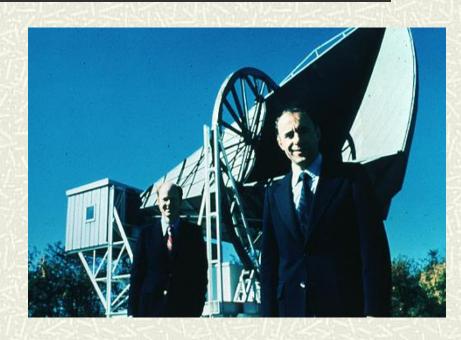


Bonus: cosmic background radiation

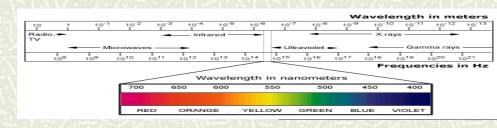
CMB discovered accidentally

- **♯** Universal signal (1965)
- **♯** Low frequency (microwave)
- **■** Low temperature (3K)

Echo of Big Bang!



Penzias and Wilson





NATURE | NEWS

Einstein's lost theory uncovered

Physicist explored the idea of a steady-state Universe in 1931.

Davide Castelyecchi

24 February 2014



New Discovery Reveals Einsteir Tried To Devise A Steady State Model Of The Universe



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Almost 20 years before the late Fred Hoyle and his colleagues devised the Steady State Theory, Albert Einstein toyed with a similar idea: that the universe was eternal, expanding outward with a consistent input of spontaneously generating matter.

An Irish physicist came across the paper last year and could hardly believe According to this week's article in Nature,

model of the universe very different to today's Big Bang Theory.





SCIENTIFIC $\mathbf{AMERICAN}^{\scriptscriptstyle{\mathsf{M}}}$



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Einstein's Lost Theory Uncovered

The famous physicist explored the idea of a steady-state universe in 1931

nature

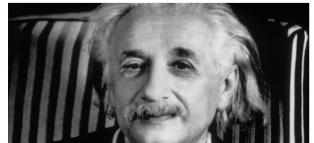
Feb 25, 2014 | By Davide Castelvecchi and Nature magazine

A manuscript that lay unnoticed by scientists for decades has revealed that Albert Einstein once dabbled with an





archive

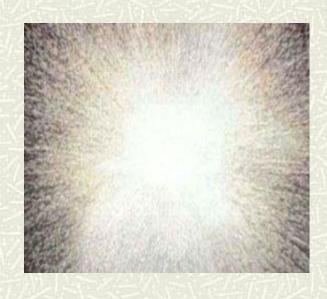




08:42 Gardaí investigate death of woman in Dublin 08:25 Flannery faces call from all parties to attend

The way back isn't so simple

The big bang – is it true?



Superhot, superdense

Expanding and cooling over time

 $\sqrt{1}$. The expansion of the U

√ 2. The abundance of H and He

 $\sqrt{3}$. The distribution of galaxies

√ 4. The cosmic background radiation



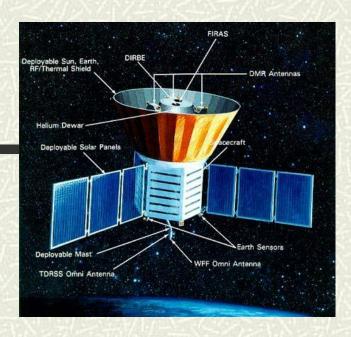
Cosmology today

- Satellite measurements of CMB

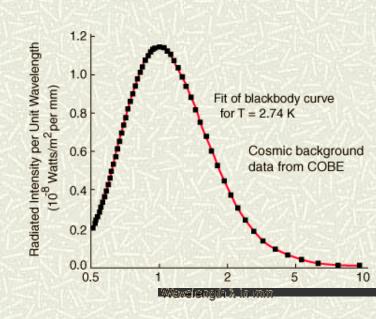
 No interference from atmosphere
- Expected temperature

 Expected frequency
- Full spectrum

 Perfect blackbody spectrum
- **Perturbations**Variation of 1 in 10⁵



COBE satellite (1992)



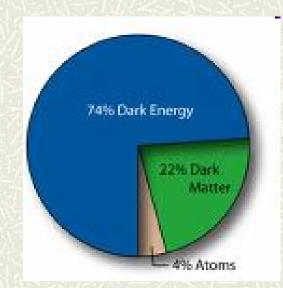
Detailed big bang model: A-CDM

A universe containing ordinary matter, dark matter and dark energy

1. Ordinary matter: 4% (astrophysics)

2. Dark matter: 22% (astrophysics)

3. Dark energy: 74% (supernova, CMB)



1-CDM

/AC-12

History of DM and DE

 $\Omega = 1$



Dark Matter

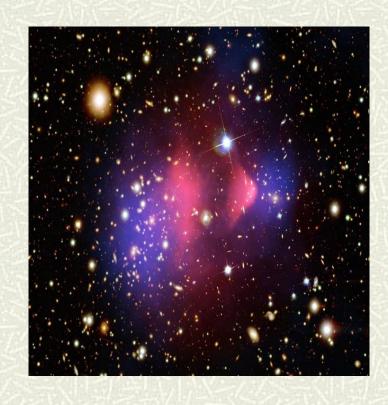
- # First suggested in 1930s
- # Stellar motion

normal gravitational effect but cannot be seen directly

- # Explains motion of stars
- # Explains motion of galaxies
- # Explains gravitational lensing

Matter =
$$OM(30\%) + DM(70\%)$$

Compatible with nucleosynthesis

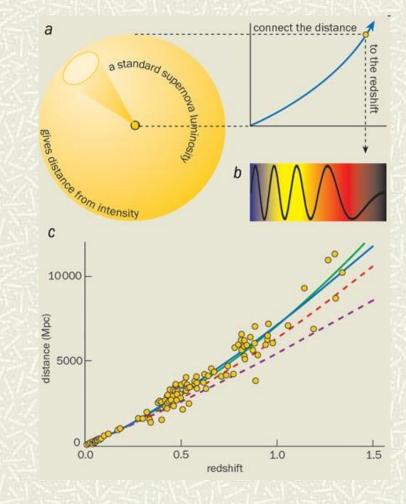


$$\Omega = 0.3$$

Dark energy

- **■** Measurements of supernovae (1998)
- # Furthest galaxies too far away
- # Cosmic expansion accelerating
- **♯** Energy of vacuum?
- **■** Not well understood





The big bang model - problems

■ Nature of dark energy?

Role in BB?



Particle experiments?

The singularity problem

What banged?

What does time zero mean?

Intersection of history with today's science





No-boundary universe

