

Revolutions in Science

A brief history of cosmology

Cormac O’Raifeartaigh FRAS

UCD School of Physics 2019

Cosmology: the study of the universe

Is the universe finite?

How big is it?

Is the universe eternal?

How old is it?

How did the universe begin?

How will it end?

What is the nature of time?

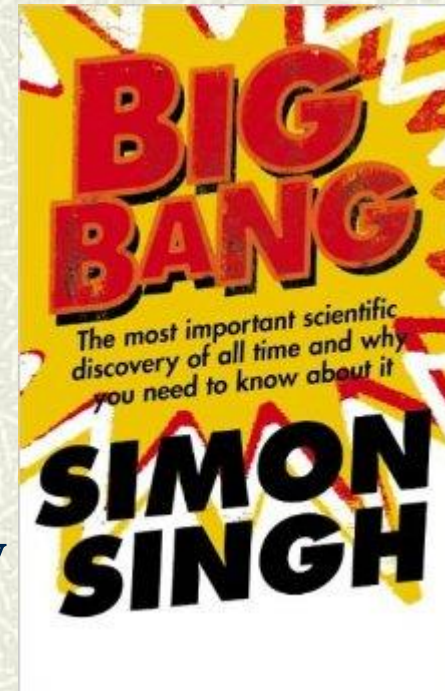
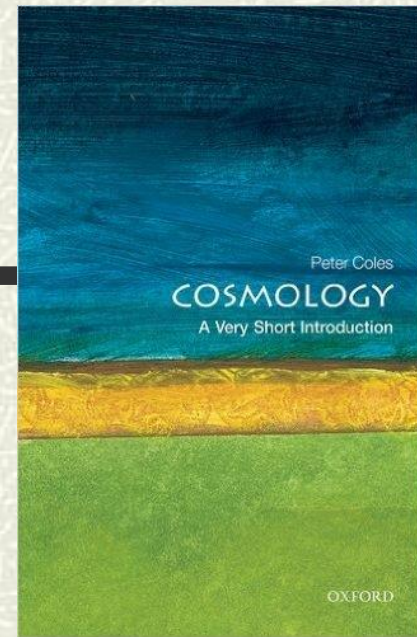
What is the nature of space?



Science or philosophy?

Useful Texts

- **A Very Short Introduction to Cosmology**
Peter Coles (OUP)
- **The Big Bang**
Simon Singh
- **Cosmology**
Edward Harrison (CUP)
- **The Cambridge Illustrated History of Astronomy**
Michael Hoskin (CUP)



Webpage: Antimatter


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Earliest cosmology

I The age of magic

Every that happened due to spirits

- **The anthropomorphic universe**

Fashioned in the image of mankind

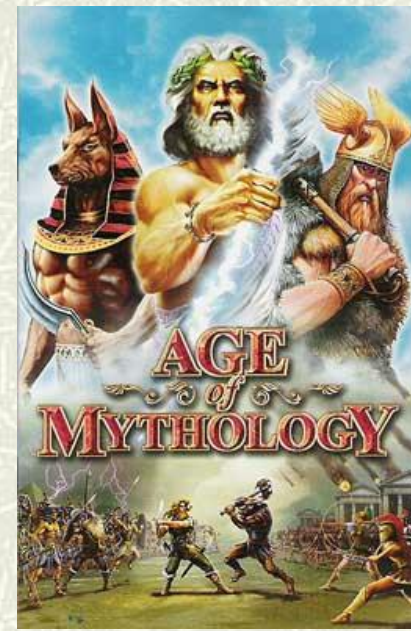
II The age of mythology

Creation myths

- **The anthropocentric universe**

Spirits become remote Gods and Goddesses

*Sumerians, Assyrians, Babylonians, Minoans,
Greek, Jews, Chinese, Norse, Celts, Mayans*



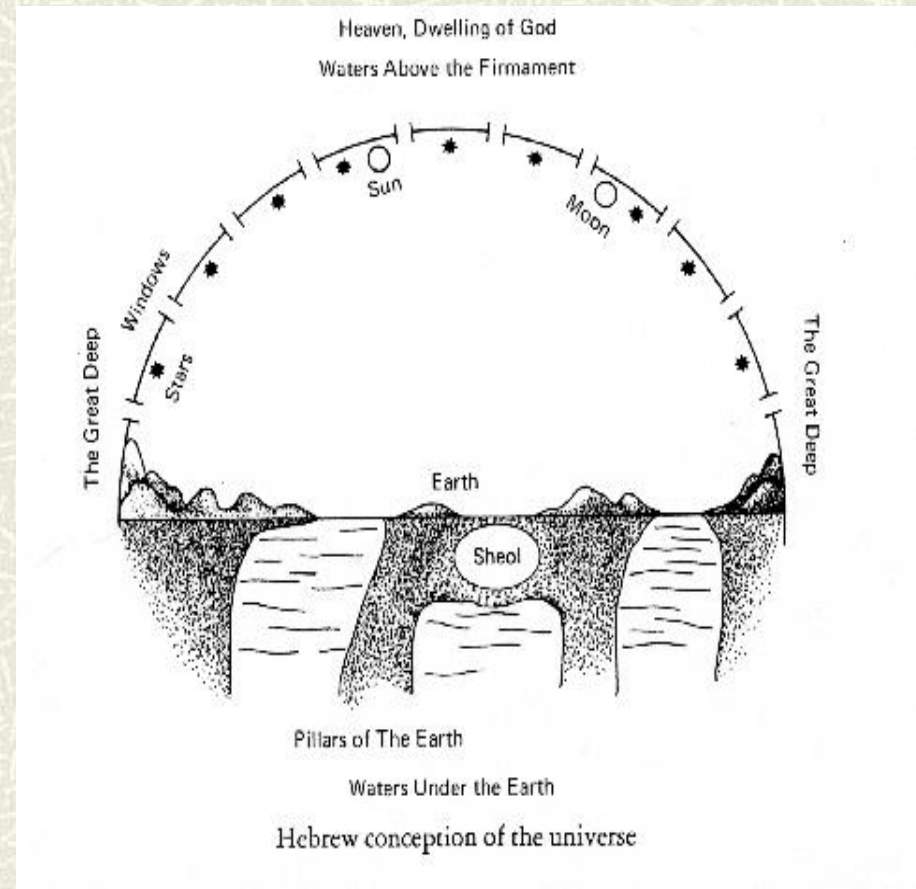
Early mythology

Babylonians, Egyptians and Hebrews (6000 BC)

- Earth in the center of an oyster, covered by dome
- Water underneath and overhead, closed in on all sides
- Moderate dimensions
- Sun, moon and stars progress across dome
- Universe guided by deity, explained by myth

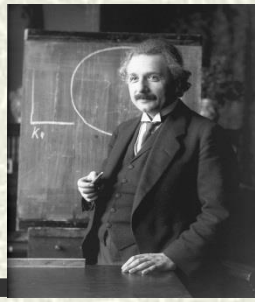


The Babylonian universe



The Hebrew universe

Earliest observations: astronomy



Chaldean priests (3800 BC)

- ⌘ Timetables of the motion of the stars and the planets
- ⌘ Stars stationary, planets move across a lane in the sky (zodiac)
- ⌘ Length of the year
- ⌘ Astronomy and astrology, calanders
- ⌘ Predicted astronomical events

Observation without explanation



The Babylonians

- **Observations of the stars**

Nabonassar (747 BC)

- **Calendar**

360 days = 1 year

Lunar months: varied

- **Measuring time**

The gnomon (sundial); 12 parts of the day

1 hour = 60 minutes: 1 minute = 60 seconds

- **Observations of the planets**

- *The Zodiac*

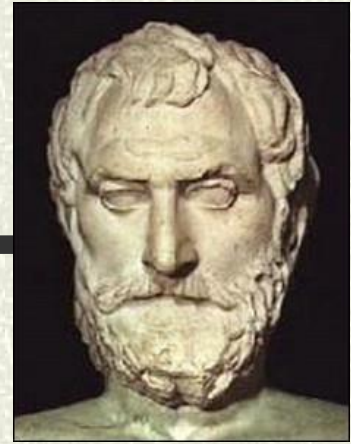
- *Circle on which the planets moved*



The cosmology of ancient Greece

Ionian philosophers (6th cent BC)

Natural causes: not concerned with deities



✚ **Thales of Miletos** (625-547 BC)

What is the basic material of the Universe? Water

Earth floats on water

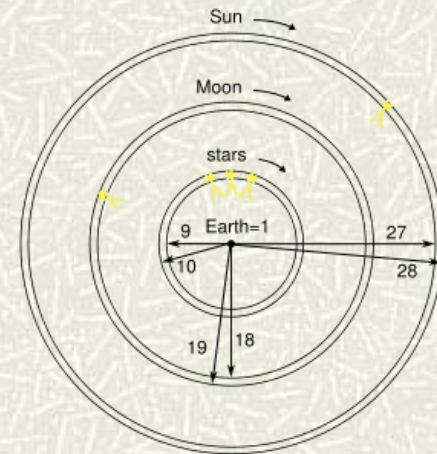
✚ **Anaximander of Miletos** (610-545 BC)

Raw material of universe: undefined apeiron

Earth floats in universe infinite in time and space

Mechanical model of the universe

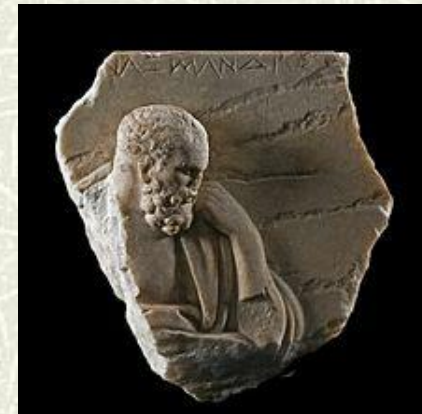
Estimated sizes and distances to sun and moon



✚ **Anaximenes of Miletos** (585-528 BC)

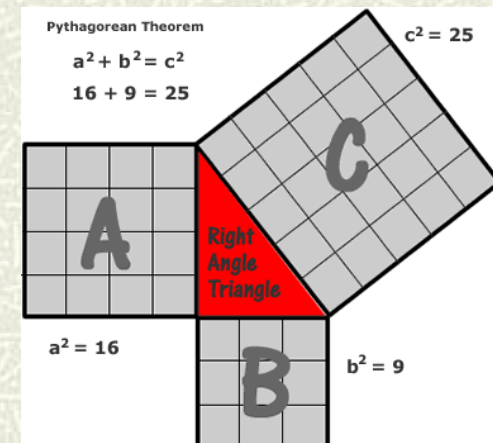
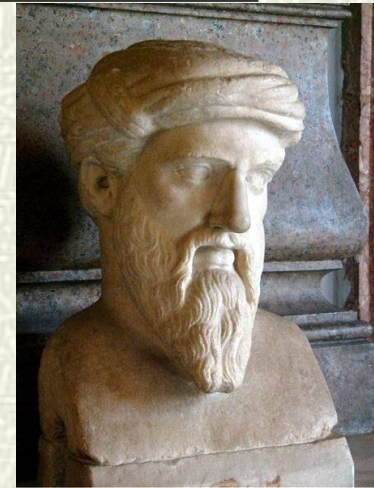
Raw material: air. Earth floats on air

Stars evolved from earth ; float on sphere above



Pythagoras of Samos (570-495 BC)

- ⌘ *Founder of science*
- ⌘ *Mathematization of experience: numbers sacred and eternal*
- ⌘ *Highest form of philosophy*
- ⌘ *All things have form, all form defined by numbers*
- ⌘ *Pitch of a note depends on the length of string that produces it*
- ⌘ *Intervals in the scale produced by simple numeric ratios*
- ⌘ *Reality could be reduced to number series and number ratios*



Theorem: *Areas of the two smaller squares of the sides of right-angled triangle will equal the area of the larger square*

The Pythagorean universe

- **Earth is a sphere**

Ships on horizon

Lunar eclipses

Shape of moon and sun

- **Orbits**

Sun, moon and planets revolve around central fire

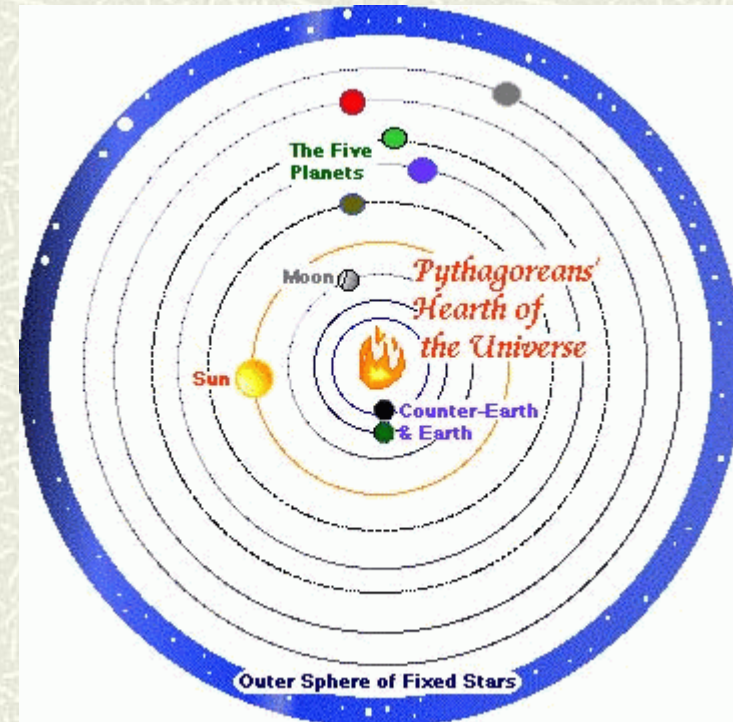
Each planet hums on a different pitch

Intervals between orbits governed by laws of harmony

- **Major influence**

Plato

Kepler



Downfall of the Pythagorans

- ✚ **Discovery of irrational numbers**

The diagonal of a square

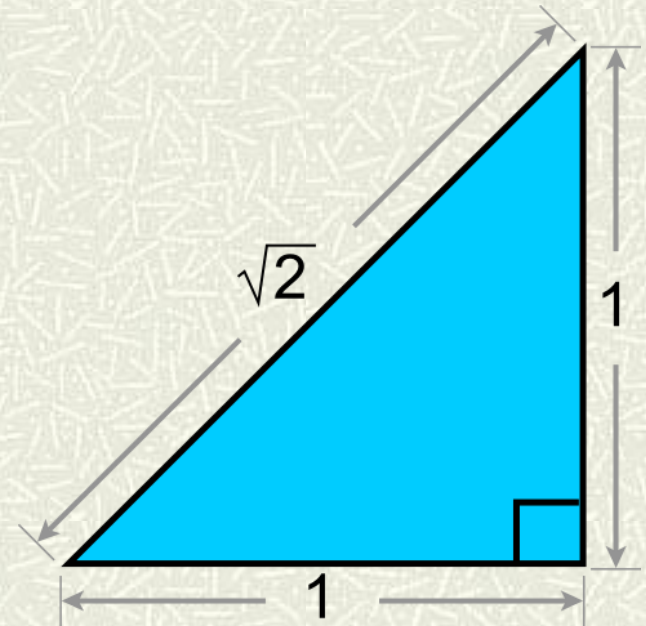
- ✚ **Arithmetic and geometry**

Breakdown in point-to-point correspondence

- ✚ **Tried to keep secret**

- ✚ **Dissolution of brotherhood**

Also due to egalitarian practices, socialist society



Pythagoras: founder of European culture, source of Platonism

Legacy of Pythagorean School

Spherical earth

Attracts everything to its center

Aristarchus of Samos (310 BC)

"On the sizes and distances of the sun and moon"

Eratosthenes

Measured circumference of earth from summer solstice

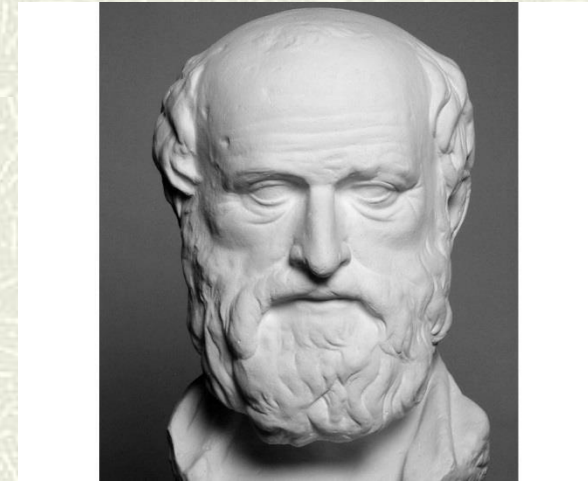
Deduced relative size of moon (from lunar eclipse)

Deduced distance of moon from earth (from geometry)

Deduced sun-earth/moon-earth distance (from half-moon)

Deduced sun-earth distance

Deduced relative size of sun (from solar eclipse)

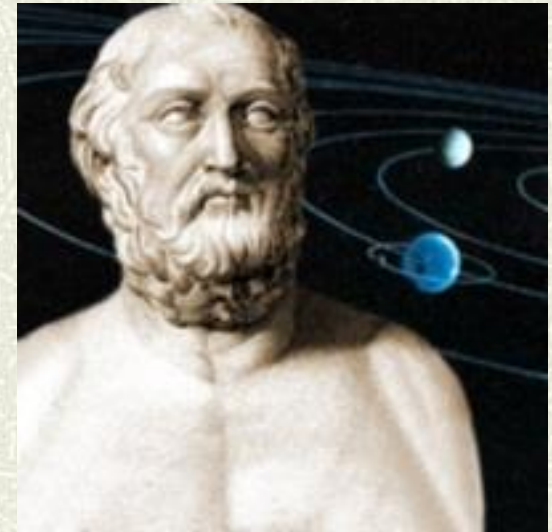


Eratosthenes 276 –195 BC

$$r_S / d_S = r_M / d_M$$

Philolaus and Aristarchus

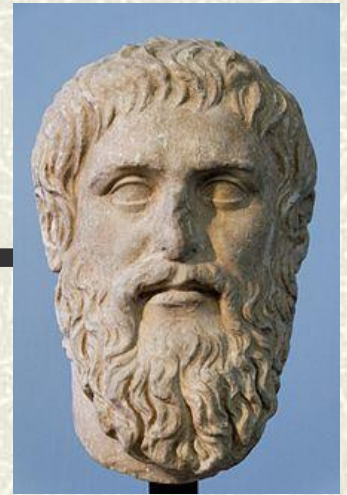
- # *Earth sphere has motion: rotates about its own axis*
- # *Daily revolution of the sky caused by earth's own motion*
- # *Separated day and night, annual motion of the planets*
- # *Earth orbits sun*
- # *First suggestion of heliocentric system*



Aristarchus 310-230 BC

Quoted by **Archimedes**: yet heliocentric system was discarded

Plato and Aristotle



- Heroic age followed by decline

Should have been Aristarchus - Copernicus - Galileo

- Plato:** certainty in mathematics, not naturalism

Shape of the world must be a perfect sphere

All motion perfect circles at uniform speed

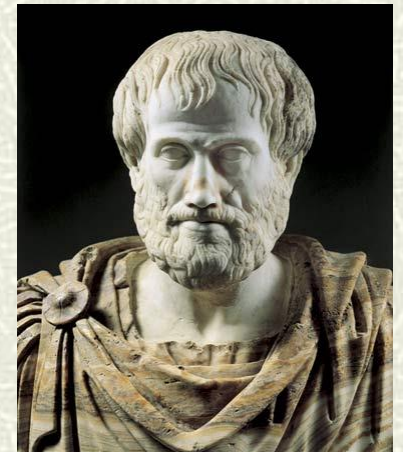
- Science dominated by **Aristotle** (logic and empiricism)

God spins the world from outside it, not from center

Earth and moon space subject to change: nowhere else

All celestial bodies orbit earth in perfect circles

Plato 424 – 348 BC



Aristotle 384 – 322 BC

Spherical universe rotates about spherical earth

The dogma of Plato and Aristotle

- ⌘ Planetary motion must be shown to be result of circular motions

Aristotle: used 54 spheres to account for motion of the planets

- ⌘ **Ptolemy** (AD 150) : ultimate earth-centered model

Complicated epicycle system for circular motion of celestial objects

Ferris Wheel universe

- ⌘ **Tradition**

Enshrined in 'He Magele Syntaxis' (later 'The Almagest')

Kept alive by Islamic scholars during the middle ages

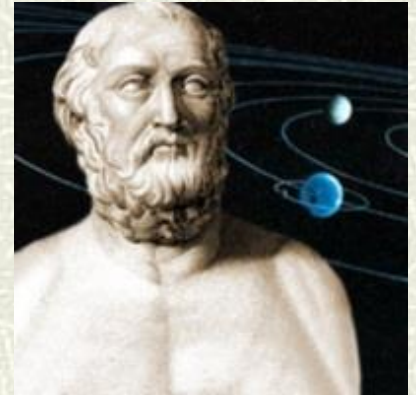
Re-introduced to Europe in 1175 - 1600

Dualism of celestial and
terrestrial motion
Immobility of earth
All heavenly motion
perfectly circular

- ⌘ **Dogma dismissed reality: 3 fundamental conceits**

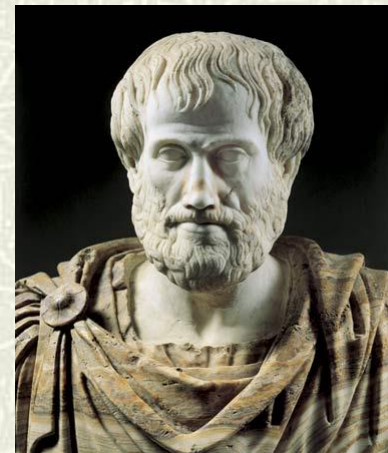
Why was Aristarchus universe discarded?

- # Objects fall towards earth
- # No wind blowing against us
- # No obvious motion of stars
(stars too far away to observe stellar parallax)
- # Heliocentric model rejected
Geocentric model retained
- # One snag: motion of planets
Anomaly in paradigm



Aristarchus 310-230 BC

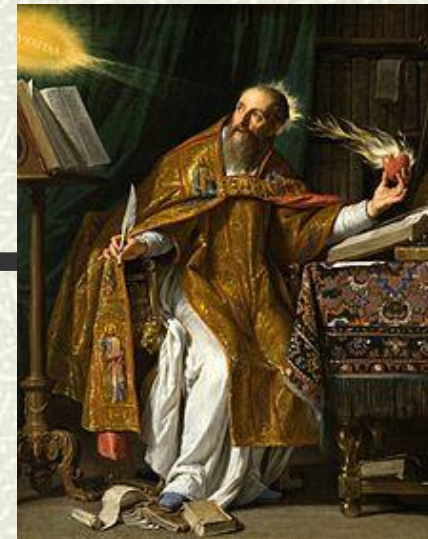
Aristotle 384 – 322 BC



Towards the Middle Ages

Platonism adopted by Christianity (**St Augustine** etc)

*Ignored early Greeks, adopted only **Plato's** philosophy: neoplatonism
11cent AD: same view as 5th cent BC*



St. Augustine (354-430)

Medieval philosophy

Aristotle's 55 spheres, Ptolemy's 40 epicycles

Replaced by 10 revolving spheres: disregarded stellar observation

1000 AD: Portolano charts for navigation

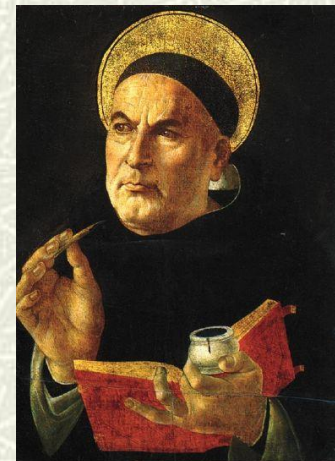
12-16th century: Aristotelean philosophy adopted

Muslims: carried fragments of Euclid, Archimedes and Aristotle to Europe

Improved Calendric astronomy and planetary tables

Important Indian numerals (including zero) and algebra

*Christianity and Aristotelianism (**Thomas Aquinas**)*



St. Aquinas (1225-1274)

Aquinas: first proof of God (based on Aristotelian physics)

Unmoved mover = God

Summary

By 1500 AD, Europe 'knew' less than Archimedes in 200 BC

1. Splitting of Universe into 2 spheres
2. Geocentric dogma
3. Uniform motion in perfect circles dogma
4. Divorce of mathematics from science
5. Divorce of experiment from science

These 5 handicaps were to be overcome by
Copernicus, Kepler and Galileo

The Renaissance



Roger Bacon (1220-1292)

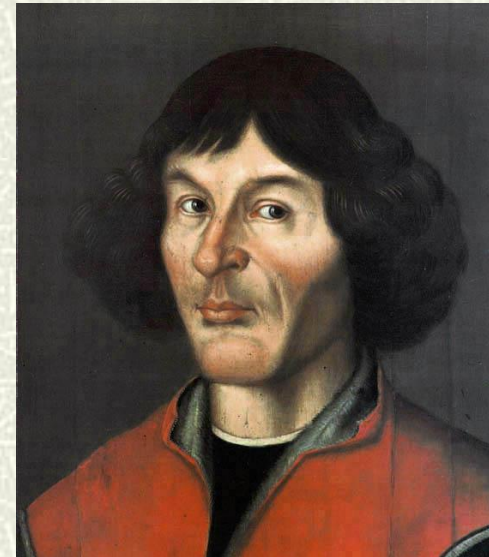


William of Ockham (1280-1349)

- **Universities: Revival of learning**
Bologna, Paris, Oxford and Cambridge
- **Roger Bacon, William of Ockham and Albert the Great**
Study of nature: empiricism
- **Ockham's razor**
Opposed Aquinas
"All other things equal, simpler explanation more likely"
- **Nicole d'Oresme**
Geocentric universe not proven
- **Nicholas of Cusa**
Earth not hub of universe
- **Nicolas Copernicus**

Nicolas Copernicus (1473-1543)

- Born in Poland (Prussia) in 1473
- Studied at the Universities of Krakow, Bologna and Padua
- Trained as Catholic canon
- Polyglot, polymath
- Simple instruments for observing the sky
- Used observations of the Chaldeans and the Greeks
- Improved astronomical tables
- Published *Commentariolus* (1514)



Nicolas Copernicus
(1473-1543)

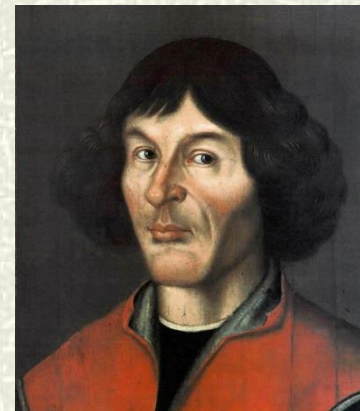
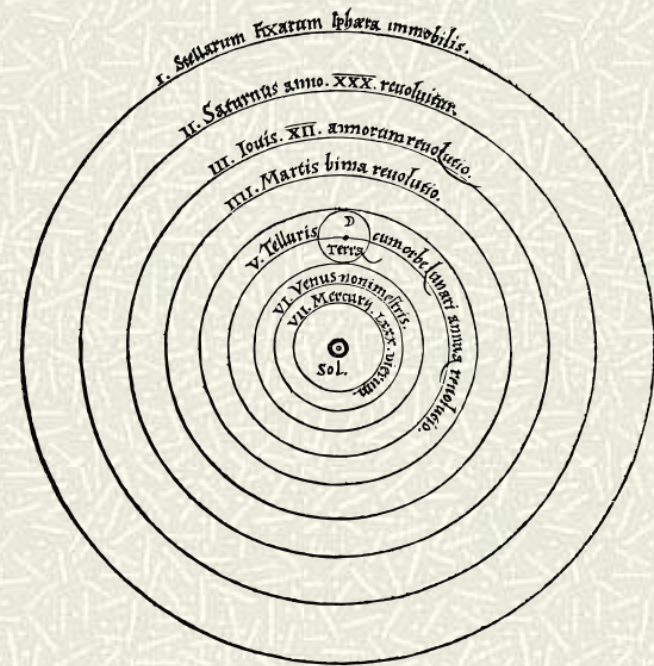
Nicolas Copernicus

Commentariolus (1514) : 7 axioms

1. Heavenly bodies do not share a common center
2. Earth is not the centre of the universe
3. Centre of the universe is near the sun
4. Earth's distance to sun negligible compared to distance to stars
5. Daily motion of stars is due to rotation of the earth
6. Earth, planets revolve around sun
7. Retrograde motion of the planets is due to motion of earth

Distributed to friends, academics: became well-known

Earth cannot be stationary, Ptolemy system inconsistent
Heliocentric model (aware of early Greeks)



Nicolas Copernicus
(1473-1543)

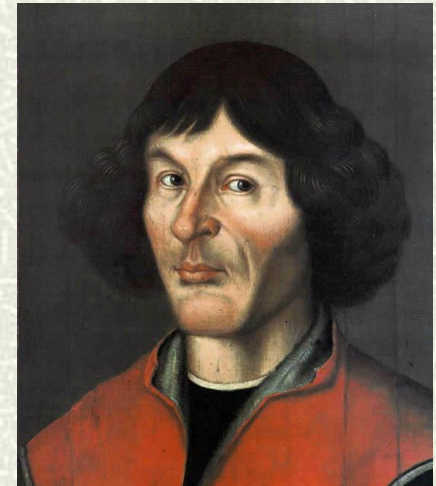
Nicolas Copernicus

Book of the Revolutions of the Heavenly Spheres (1543)

- Published with aid of Georg von Lauchen (Rheticus)
- Technical version of Commentariolus
- Ahead of its time (earth just a planet like others)

Problems

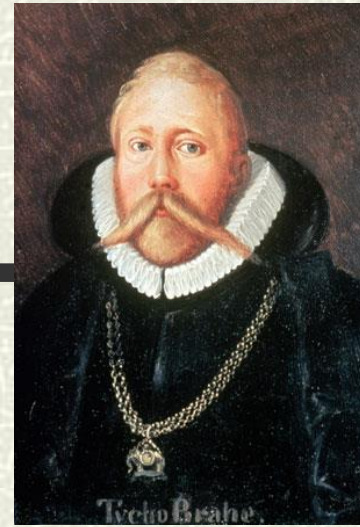
- Copernican system quite complicated (48 epicycles)
- Less accurate predictions of planetary motion
- Universe de-centralized: no natural center of directions
- Undermined by preface (Osiander)



Nicolas Copernicus
(1473-1543)

Not adopted for many years: unnecessary, opposed by Church

2. Tycho de Brahe



Tycho de Brahe
1546-1601

Danish nobleman

Observational astronomy to new level

Own observatory on island of Hven

Expensive instruments: precise and continuous data

Observed distant nova (SN1572)

Observed distant comet (1577)

Heavens not immutable

Sympathetic to Copernican ideas

‘Mixed’ model of universe (1588)

Migrated to Prague (imperial mathematician)

Employed *Kepler* to study Mars

Gave data sparingly to Kepler

Died from drink (1600)



The Tychonian model



Tycho de Brahe
1546-1601

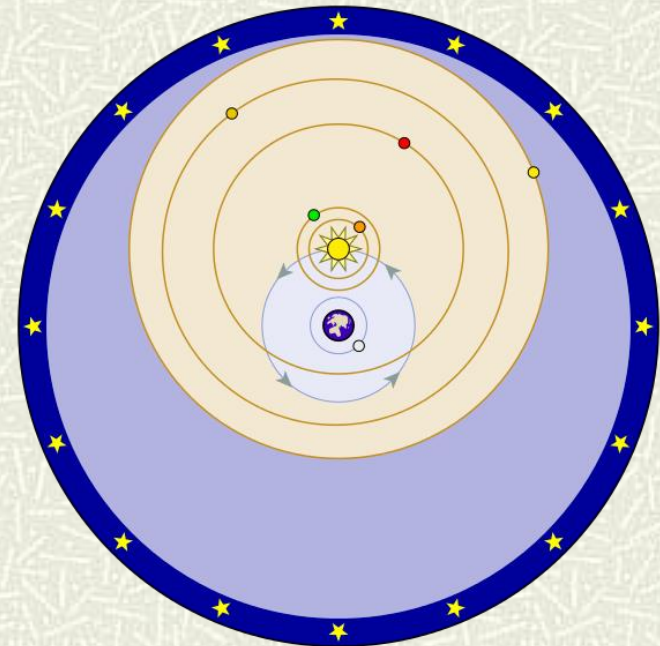
Observational astronomy to new level (own observatory)
Observations showed heavens not immutable

Sympathetic to Copernican ideas
Critical of Copernicus's observations

Convinced of immobility of earth
Lack of stellar parallax v. sizes of stars
Aristotlean physics, religion

‘Mixed’ model of universe (1588)
Planets orbit sun
Sun and moon orbit earth

Geo-heliocentric model



Johannes Kepler

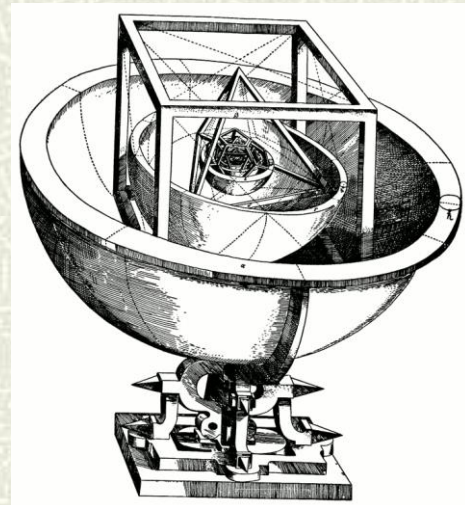
Born in Germany: not high-born
Mathematical prodigy
Attended University of Tübingen
Trained as Lutheran cleric
Became teacher of mathematics and astronomy in Graz

Defended **Copernicus** while still a young priest
Mysterium Cosmographicum (1596)
Reconciled Copernican model with scripture
First public commitment by astronomer to Copernican system
Platonic solid model of the cosmos

Hired by Tycho de Brahe (1600)
Given Mars project, but limited data
Not convinced by Tychonian model
Tycho dies after a few months



Johannes Kepler
1571-1630



Platonic solids

Johannes Kepler



1571-1630

Death of Tycho de Brahe

Kepler becomes Imperial Mathematicus (1601)

Access to all of Tycho's data

Solves problem of the orbit of Mars (8 years)

Mars orbits sun in elliptical path

All planets move in elliptical paths: no epicycles

First accurate heliocentric model of universe

"Astronomia Nova " (1609)

Simple, elegant model

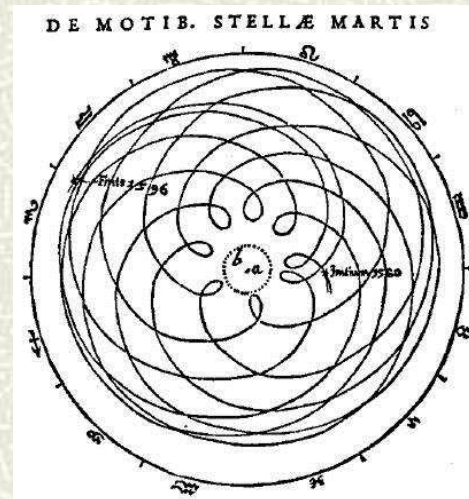
Natural laws: precise, verifiable statements based on data

Material bodies acted upon by forces

Accurate predictions

Ignored by Germans and Italians (incl. **Galileo?**)

Accepted by British (Thomas Harriot, Jeremiah Horrocks)



Kepler's universe

Kepler's planetary laws

1. Planets move in elliptical orbits (not circles)
2. Planets continuously vary their speed (not uniform)
3. Sun is at focus of ellipse (not center of circle)

Kepler's other work: founded instrumental optics

Optics (1604)

Study of refraction by lenses

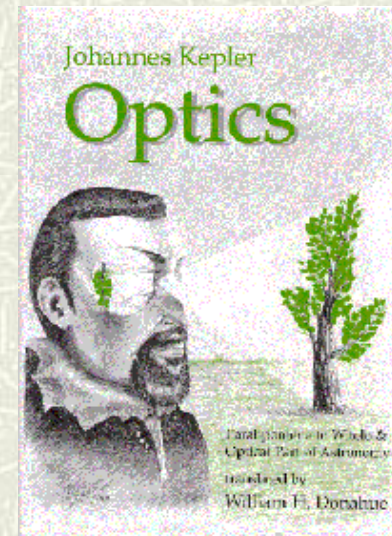
Light intensity $\propto 1/r^2$

Principle of sight, spectacles

Camera obscura



Johannes Kepler
1571-1630



Galileo Galilei



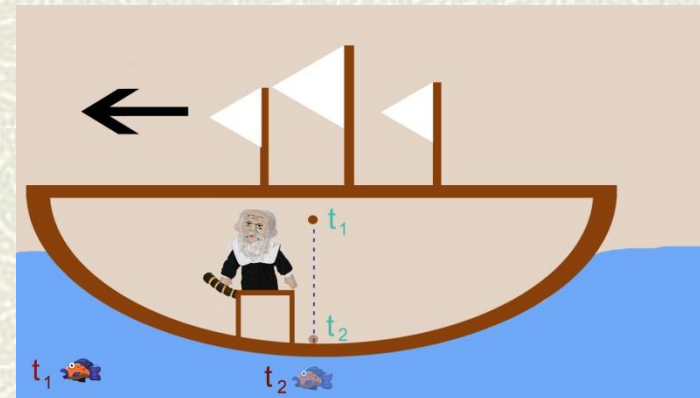
Galileo Galilei
1564-1642

Studied pendulum as student
(period depends only on length)
Invented pendulum timing device

Appointed lecturer in mathematics at the University of Pisa (1589)
Chair of Mathematics at the University of Padua (1592)

Studied mechanics in detail
Laid foundations of dynamics
Laws of motion of falling bodies
Optics: lenses, mirrors and camera obscura

Relativity of motion
Ball drops independent of ship's motion
Could earth be moving?



Relativity of motion

Galileo's astronomy



Galileo Galilei
1564-1642

Improved Dutch model of telescope (mag x 30)

Discovered Lunar surface not smooth (contradicts Ptolemy)

Discovered moons of Jupiter : earth not center of all things

Discovered phases of Venus (as predicted by Copernicus)

Discovered innumerable stars in Milky Way

Discovered sunspots: sun subject to decay etc

Defended Copernican system *Sidereus Nuncius* 1610



Measurements confirmed by Jesuits

The Galileo Affair (I)

University Aristotelians opposed Galileo

Letter to Castelli/ Grand Duchess Christina

Defended Copernican system, suggested scriptures not literal

Church Charges brought against Galileo in 1615, dropped

Church demanded proof of Copernican system

Galileo refused (couldn't prove): ignored Kepler's findings

Galileo censured by decree (1616): Copernican model outlawed



Galileo Galilei
1564-1642

The Galileo Affair (II)

1623: New pope friendly towards Galileo

1632: "*Dialogue on the Great World Systems*"

Propounds Copernican system: pope cast as Simplicio

Pope displeased, felt deceived

1633: Galileo interrogated by Inquisition

Found guilty of defying 1616 decree

Sentence: House arrest for life: *Dialogue* prohibited

After trial

Magnum Opus – dynamics

Dialogue Concerning Two New Sciences (1636)



Dialogue of the two world systems



Isaac Newton

Born in Lincolnshire 1642

Promising student

Studied mathematics at Cambridge (TC)

Studies interrupted by Great Plague of 1666

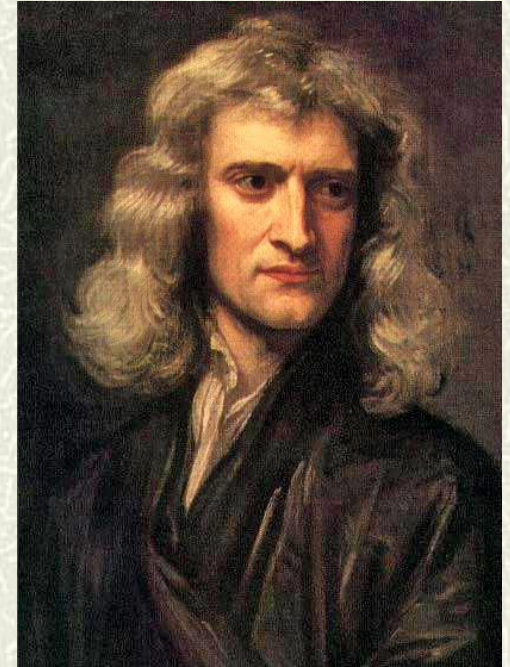
Significant work at home in calculus, optics, gravitation

Gravitation theory synthesized what went before

1. Kepler's laws (planetary motion)
2. Galileo's laws (motion of bodies on earth)

What is nature of the force that drives the planets?

What will a body do if left alone?



Isaac Newton
1642 – 1727

Key step: *Identified Keplerian orbit of moon
with Galilean orbit of projectile*

Isaac Newton (1642 – 1727)

Newton's Principia (1687):

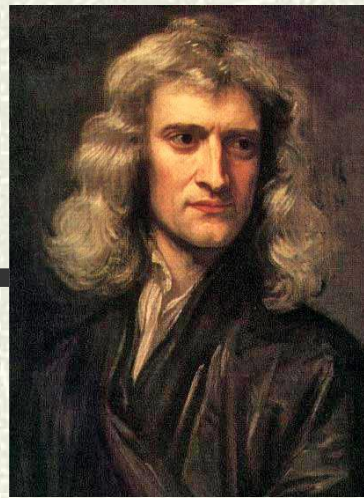
Law 1 Every body continues in its state of rest, or uniform motion in a straight line, unless compelled to change by forces acting on it

Law 2: The change in motion is proportional to the force impressed, and in the direction of the force

Law 3: To every action there is always an equal and opposite reaction: the mutual actions of two bodies upon each other are always equal, and directed to contrary parts

Laws 6 and 7: Universal Law of Gravitation $F_G = Gm_1m_2/r^2$

Isaac Newton



Isaac Newton
1642 – 1727

Universal law of gravity (1666)

Force of attraction proportional to the masses,
Inversely proportional to square of separation

$$F_G = GMm/r^2$$

Dropped for 20 years

Developed mechanics and calculus

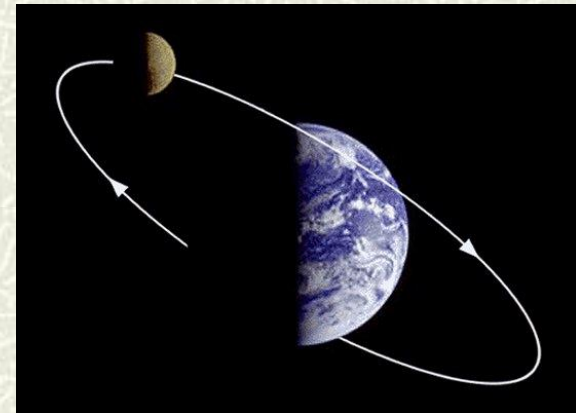
Goaded back to problem by Halley

Computes force of earth's attraction for the moon (1686)

Explains observed motion: repeated for sun

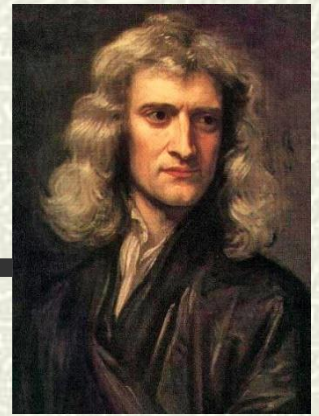
Showed orbit produced by law was Kepler ellipse

Kepler's laws arise as consequence



Moon = falling body

Newton's law of gravity



$$F_G = GMm/r^2 \quad G = \text{constant } (6.6 \times 10^{-11})$$

Unites terrestrial and celestial gravity

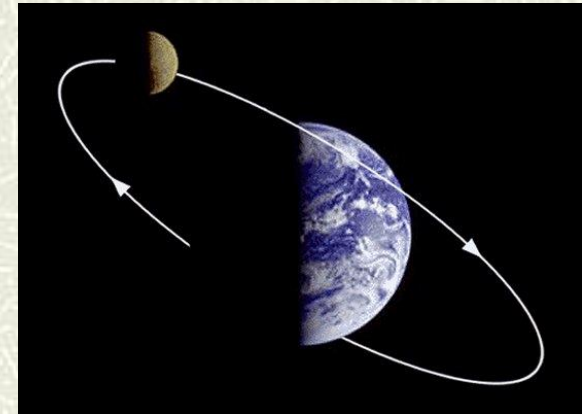
Force on apple = force on planets

F_G v. weak force (G extremely small)

F_G only seen when one mass is a planet

F_G always attractive

M, m = inertial mass (see N2)



*How does F_G act instantaneously
across huge distances?*

$$F = \frac{Gm_1m_2}{r^2}$$

The Newtonian universe



1. Force of gravity –

Attractive, weak, infinite range

2. Eternal universe (infinite in time)

No beginning, no end

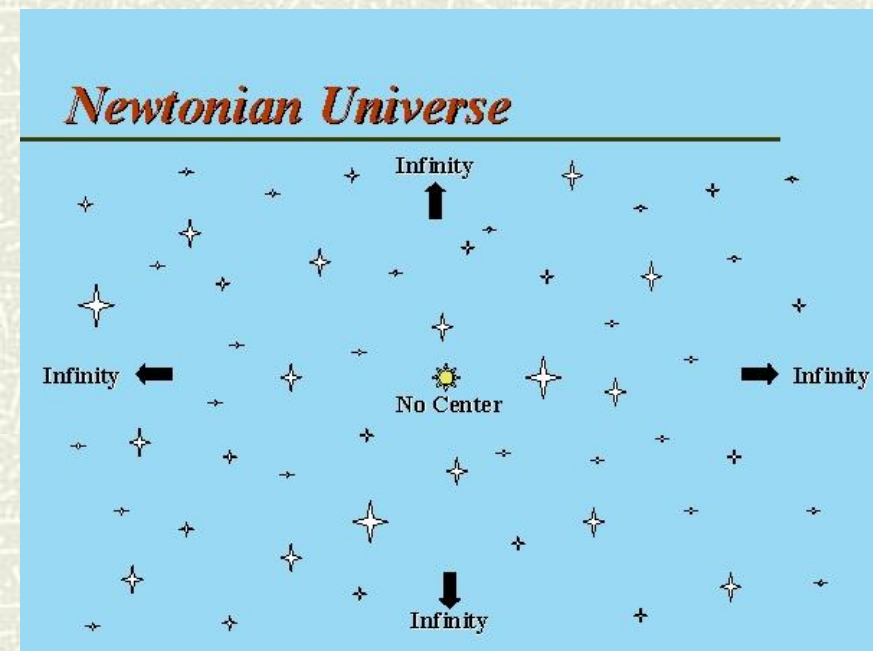
3. Bentley's paradox

Why doesn't gravity crush universe?

Finite mass in infinite space (Newton)

4. Olbers' paradox

Why is the sky dark at night?



Clockwork universe: space and time a fixed stage

Astronomy in the 19th century

1. Improved telescopes

Distances to stars

2. Stellar parallax

Bessel 1838

3. Photography

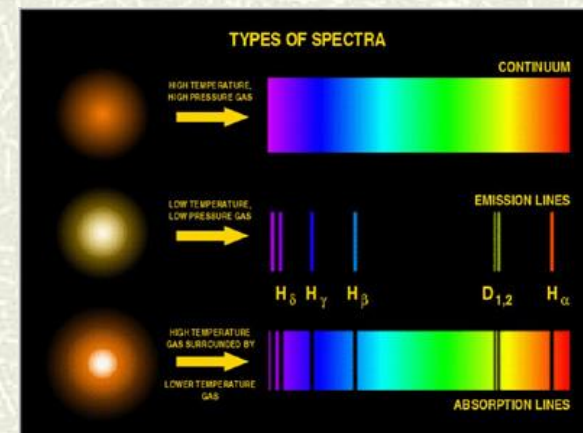
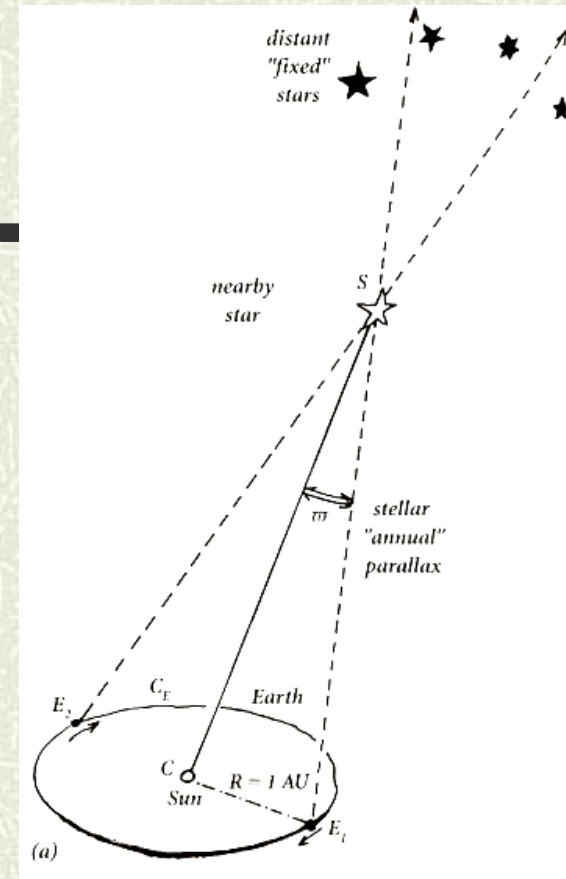
Improved images

4. Spectroscopy

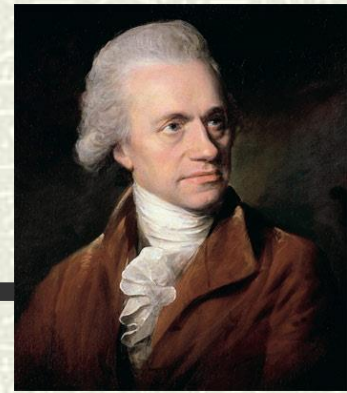
Analysis of starlight

Spectral lines of known elements

Stars made of familiar elements



The starry nebulae



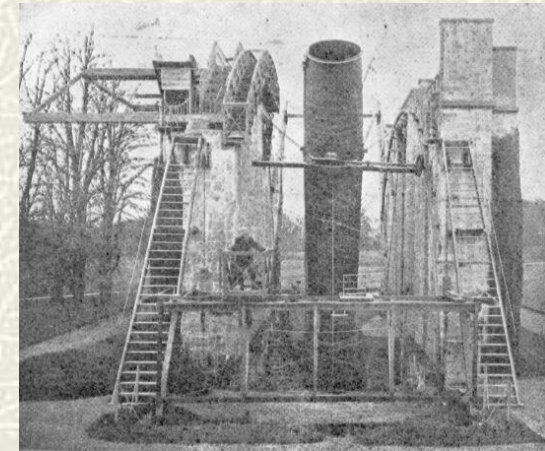
William Herschel 1738-1822

Extremely distant, cloud-like formations

Island universes? Kant, Laplace

Study of the nebulae using large telescopes

Herschel: 36-inch reflecting telescope



Some nebulae have spiral structure

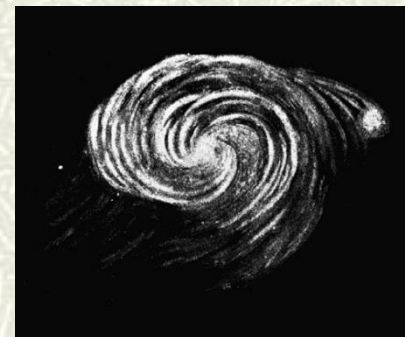
Earl of Rosse (1845): 72-inch telescope, Birr castle

Stars within spirals

Earl of Rosse: Birr castle

Distance to the nebulae?

Too distant for parallax method



William Parsons 1738-1822

The motion of the nebulae



V.M. Slipher
1875-1969

Vesto Slipher

Use of spectrograph to study nebulae

Lens speed of camera: Exposure time

First study of the nebulae (Lowell Observatory)

Light significantly blue-shifted (1912)

Doppler shift due to radial motion?

Approaching at 300 km/s

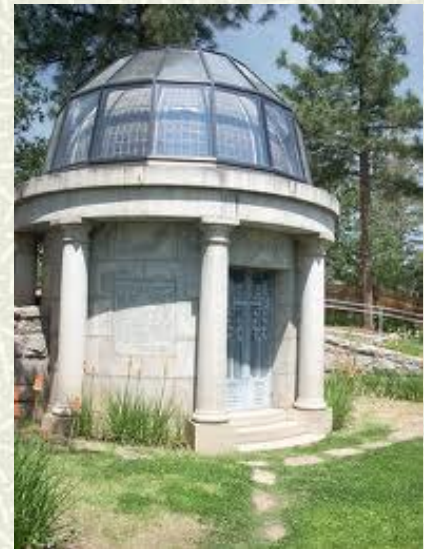
Redshift of the spiral nebulae

45 Doppler shifts (1917): all red-shifted except 4

Radial velocity outward? Large velocities: 300 to 1100 km/s

Conclusion

Distant nebulae moving away: most distant moving fastest



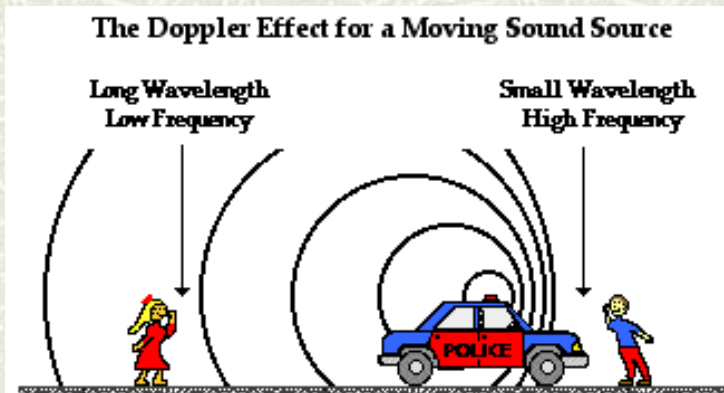
The Doppler effect

Perceived frequency of light and sound depends on relative motion of observers

Measure motion of object by measuring shift in frequency of light emitted



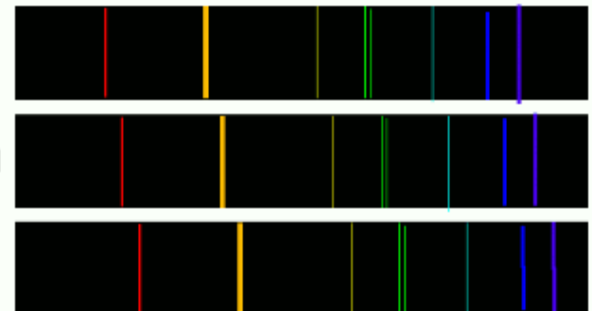
Christian Doppler
1803-1853



red shift

no motion

blue shift



The great debate



Henrietta Leavitt
1868-1921

A new measurement of distance (Leavitt)

Pulsating stars: Cepheid variables (1912)

Intrinsic luminosity related to period of pulsation

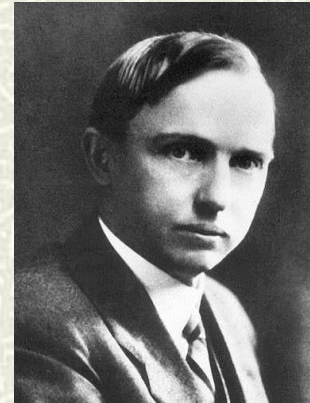
Measure distance by measuring period of luminosity

Size of Milky Way (Shapley)

Use Leavitt's law to measure Milky Way (0.5 M LY)

V bright nova in Andromeda: within Milky Way?

Harlow Shapley
1885-1972



Nebulae outside Milky Way? (Curtis)

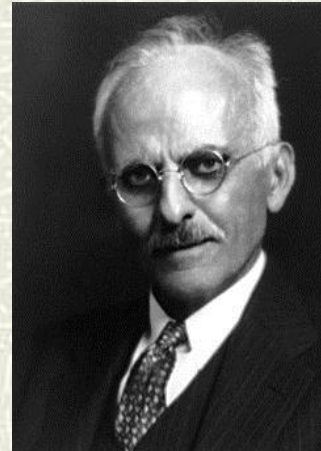
Emission lines Doppler-shifted : Vesto Slipher

Motion too great to be confined to Milky Way

Measurement of distance to nebulae (Hubble)

Well outside Milky Way

Heber Curtis
1872-1942



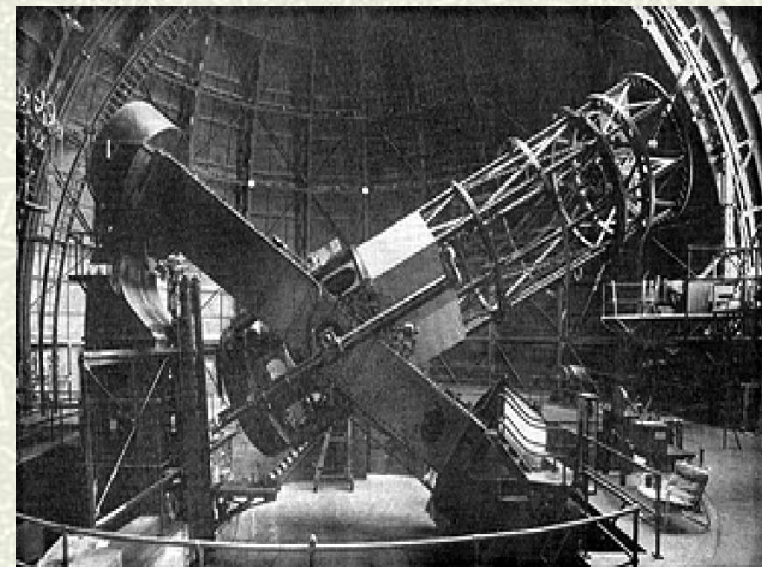
Hubble's answer (1925)

- # Edwin Hubble (1921)
- # Ambitious astronomer
- # Hooker telescope (Mt Wilson, 1917)
- # 100-inch reflector
- # Resolved Cepheid stars in nebulae
- # Known luminosity and distance
- # Far beyond Milky Way! (1925)

Nebulae are distinct galaxies



Edwin Hubble (1889-1953)



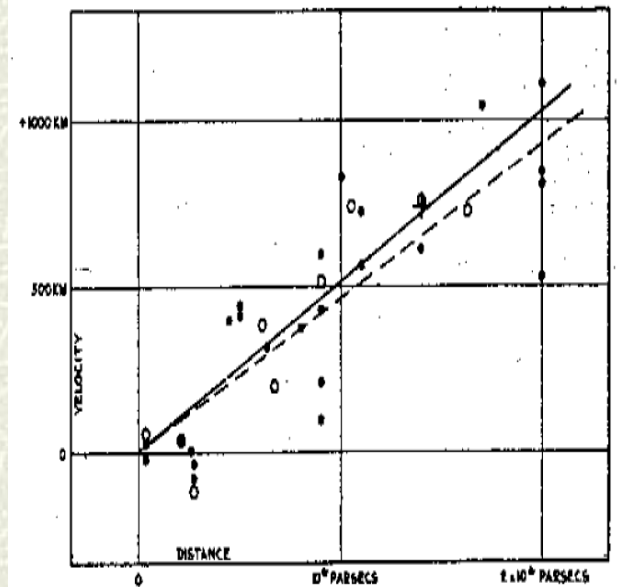
Hubble's Law (1929)

- What do the velocities of the galaxies mean?
- Is there a relation between distance and velocity ?
- Combine 24 distances with Slipher redshifts
- Approx linear relation: Hubble's law

Furthest galaxies receding fastest

Slipher not acknowledged

Velocity-Distance Relation among Extra-Galactic Nebulae.



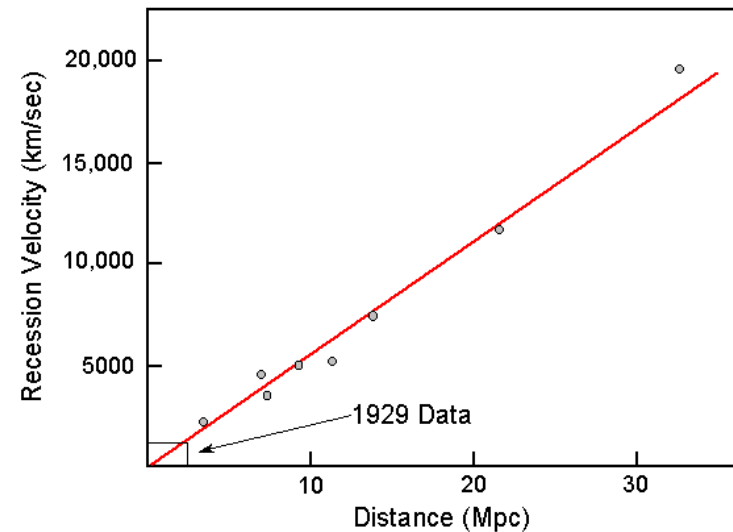
$$H = 585 \text{ km/s/Mpc}$$

Justification (1931)

- Distances for 40 galaxies
- Redshifts for 40 galaxies
- Reduced scatter – linear relation
- Justification

Explanation?

Hubble & Humason (1931)



The expanding universe

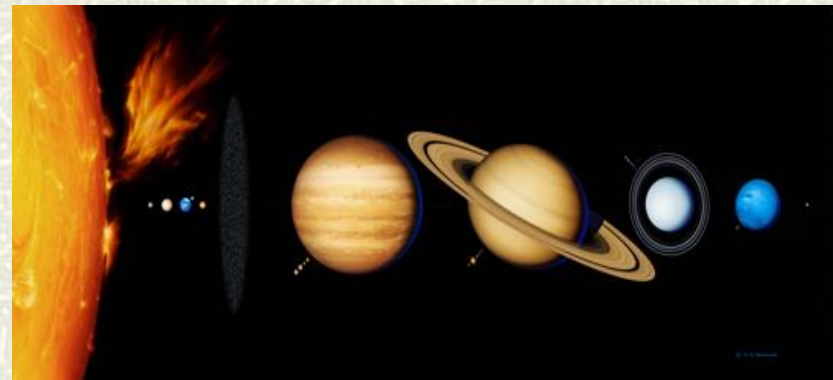
- What do the redshifts represent?
- Recession velocities for distant galaxies?
- If so, why?
- Newtonian gravity pulls in
- What is pushing out?

$$F = GMm/r^2$$

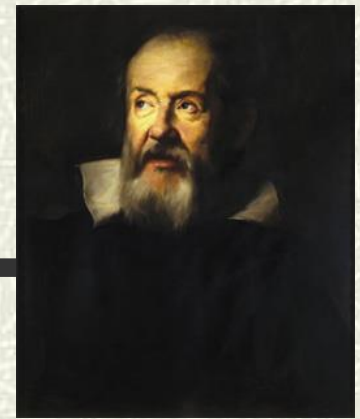


Isaac Newton

Space, time fixed



Relativity



Galileo Galilei (1564-1642)

The principle of relativity

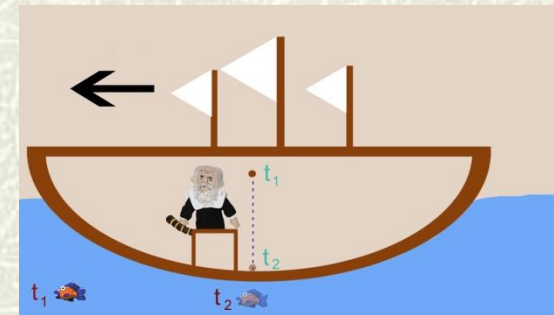
Laws of mechanics identical for observers in uniform motion

Non-accelerated motion

Galileo

Motion of ball in cabin of sailing ship

Impossible to deduce motion of ship



Application

Elizabeth I and the Irish Chieftains

Everyday experience

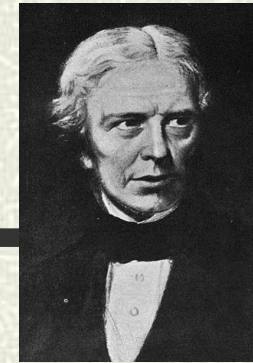
Cup of tea on train

Life on earth

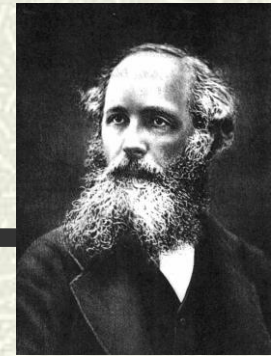


Elizabeth I (1558-1603)

Relativity in the 19th century



Michael Faraday



JC Maxwell

Electromagnetism

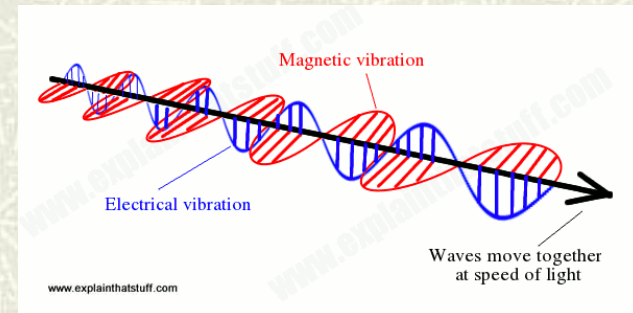
Electricity and magnetism = electromagnetism

Speed of electromagnetic wave = speed of light

Light = an electromagnetic wave

Travelling wave

Changing electric and magnetic fields



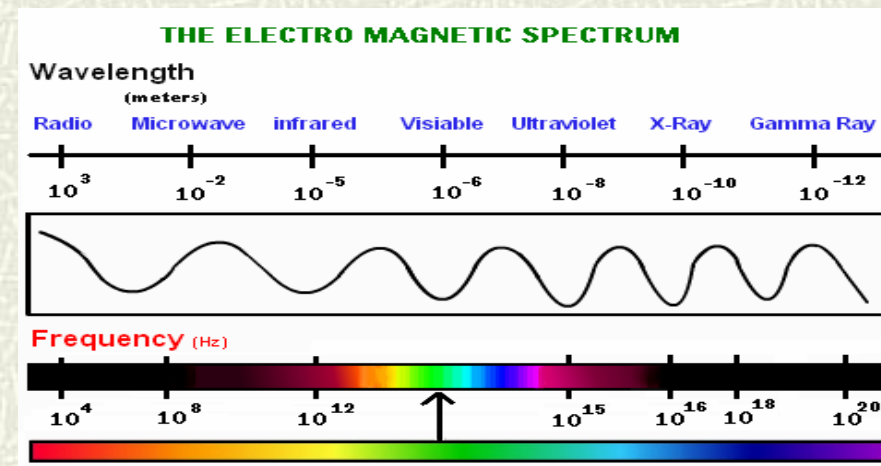
The electromagnetic spectrum

From radio waves to X-rays

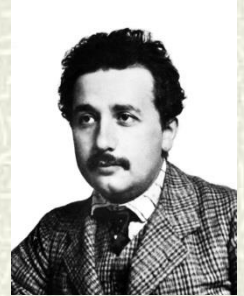
Speed of light absolute?

Fixed for all observers?

Michelson-Morley experiment



Einstein's special theory of relativity



Two new principles (1905)

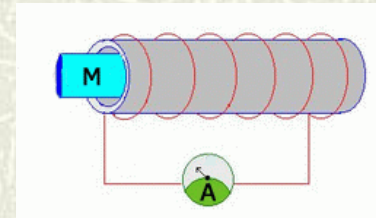
Laws of physics the same for observers in uniform motion

Speed of light the same for observers in uniform motion

Implications

Distance and time not absolute $v = s/t$

Experienced differently by bodies in motion



Predictions for high-speed bodies

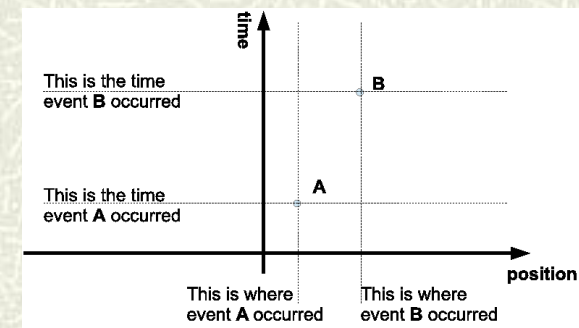
Length contraction; time dilation

Mass increase; equivalence of mass and energy $E = mc^2$

$$ds^2 = dx^2 + dy^2 + dz^2 - c^2 dt^2$$

Space + time = spacetime

Space-time invariant (Minkowski)

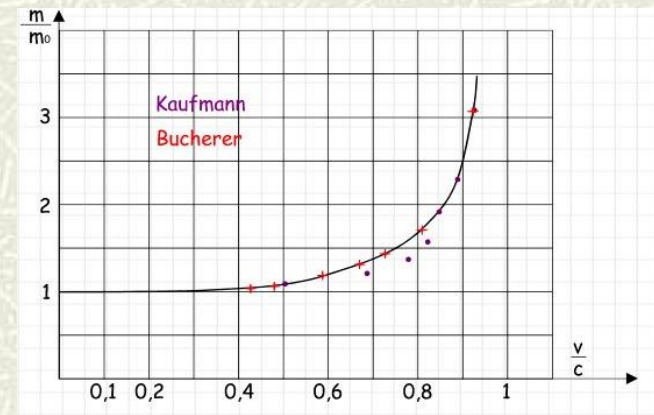


Evidence for special relativity

⌘ Mass increase

The experiments of Kaufmann and Bucherer

$$m' = \frac{m_0}{\sqrt{1 - v^2 / c^2}}$$

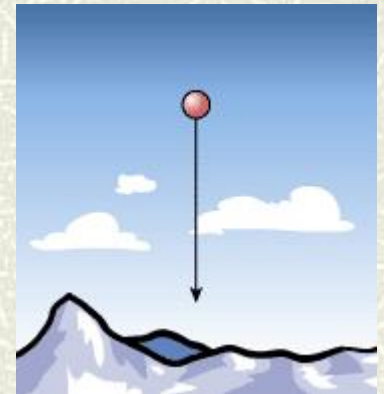


⌘ Time dilation

The long-lived muon

$$2 \mu s \rightarrow 22 \mu s$$

$$t' = \frac{t_0}{\sqrt{1 - v^2 / c^2}}$$



⌘ Invariance of the speed of light

Always measured as c

⌘ Particle experiments at the LHC

Maximum velocity = c

Mass increase

Particle creation

$$E = mc^2$$



The general theory of relativity (1916)

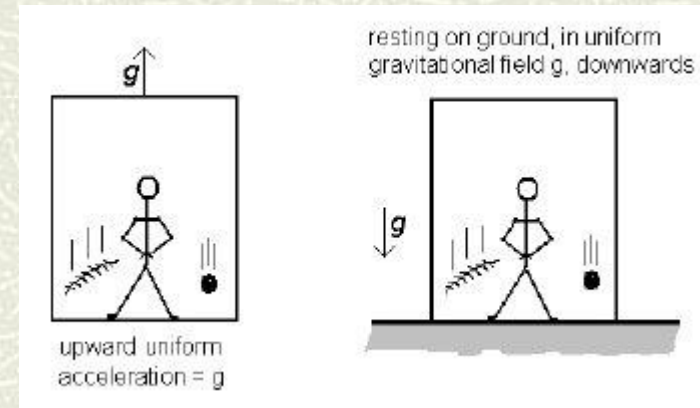
✚ Extending the special theory (1907-)

Relativity and accelerated motion?

Relativity and gravity?

✚ The principle of equivalence

Equivalence of gravity and acceleration



✚ Mach's principle

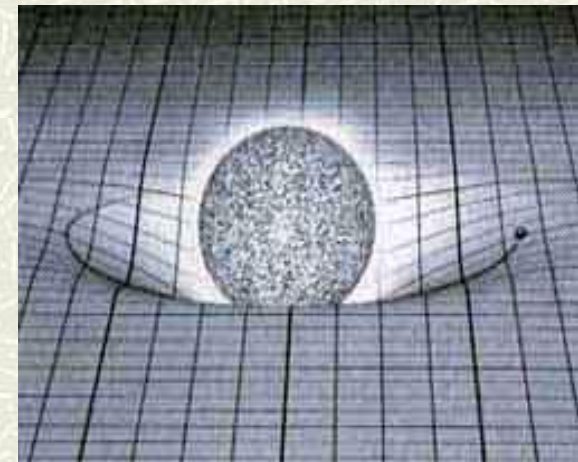
Inertial mass defined relative to matter

Space and time defined by matter

✚ A long road (1907-1915)

Curvilinear geometry, tensor algebra

Gravity = curvature of space-time



The field equations of GR (1915)



$$G_{\mu\nu} = -\kappa T_{\mu\nu}$$



10 non-linear differential equations that relate the geometry of spacetime to the density and flow of energy and momentum

SR

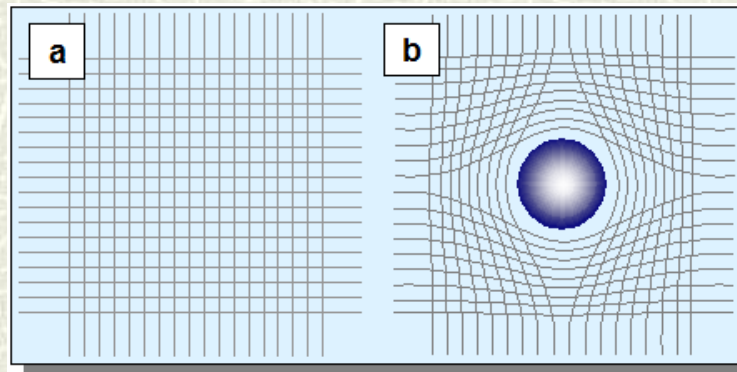
$$ds^2 = dx^2 + dy^2 + dz^2 - c^2 dt^2$$

GR

$$ds^2 = g_{\mu\nu} dx^\mu dx^\nu$$

$$ds^2 = \sum_{\mu,\nu=1}^4 n_{\mu\nu} dx^\mu dx^\nu$$

$$n_{\mu\nu} = \begin{bmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & -1 \end{bmatrix}$$



$$ds^2 = \sum_{\mu,\nu=1}^4 g_{\mu\nu} dx^\mu dx^\nu$$

$g_{\mu\nu}$: variables

Three astronomical tests (Einstein, 1916)

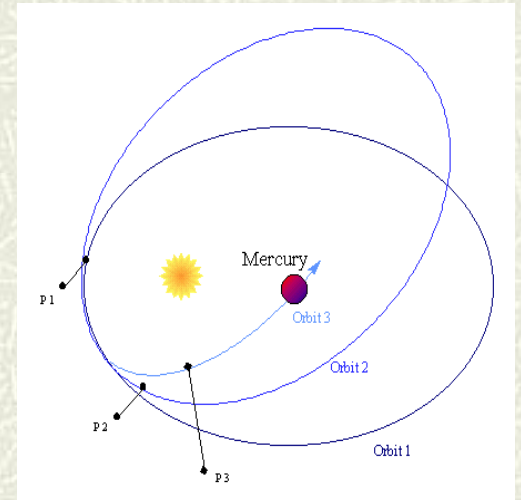
Different in principle from Newton's gravity

Small deviations in practice (weak scale)

The perihelion of Mercury

Well-known anomaly in Mercury's orbit (43" per century)

Postdicted by GR (1916)



The bending of starlight by the sun (1.7")

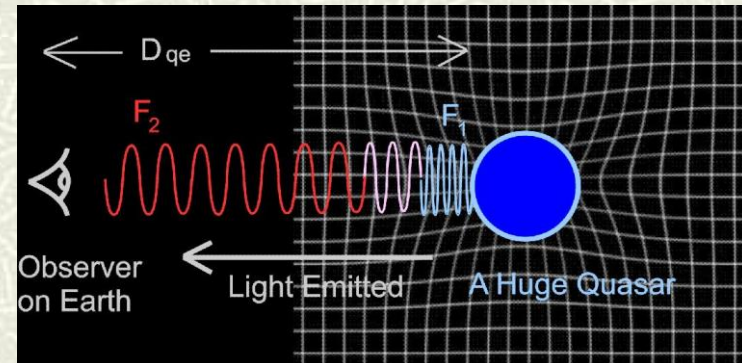
Eclipse expeditions of Eddington and Dyson (1919)

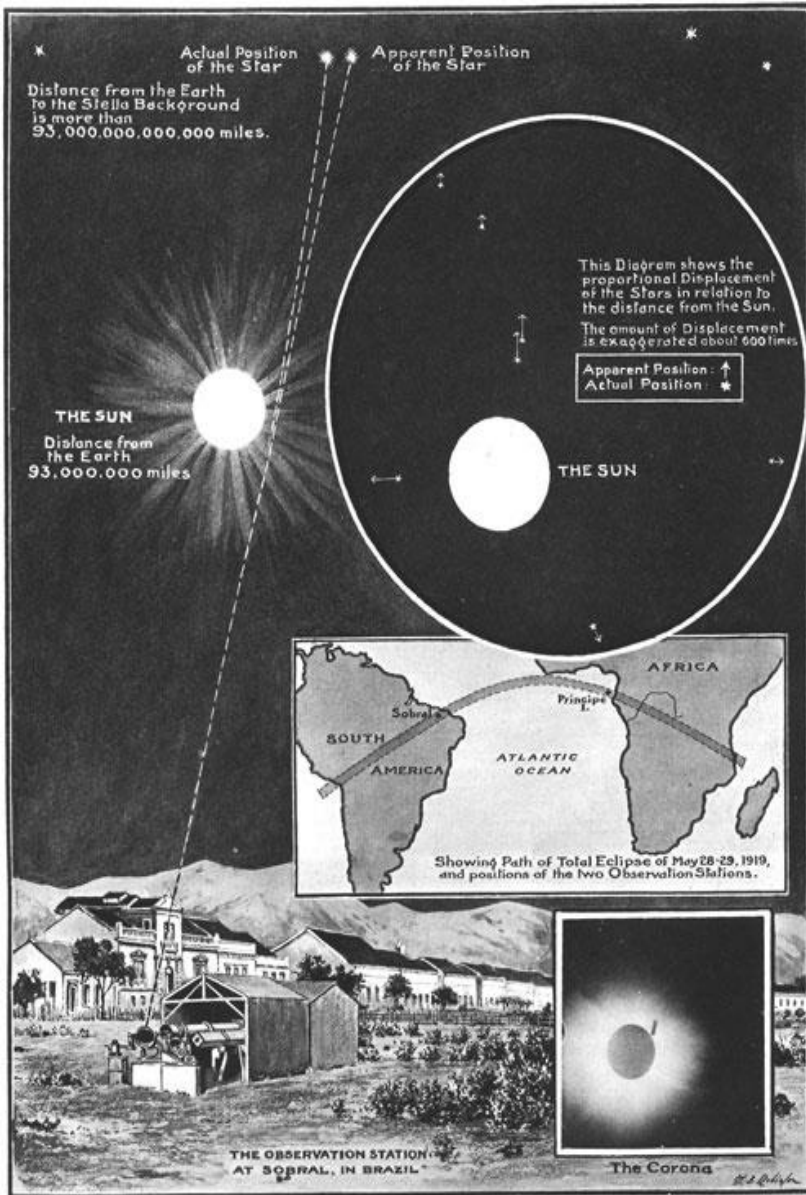
Successful measurement (large error margin)

Gravitational redshift

Time dilation in strong gravitational field

Light from a massive star redshifted?





Eclipse Results (1919)

Sobral: $1.98'' \pm 0.16$

Principe: $1.7'' \pm 0.4$

Einstein famous (1919)

LIGHTS ALL ASKEW IN THE HEAVENS

Men of Science More or Less
Agog Over Results of Eclipse
Observations.

EINSTEIN THEORY TRIUMPHS

Stars Not Where They Seemed
or Were Calculated to be,
but Nobody Need Worry.

A BOOK FOR 12 WISE MEN

No More in All the World Could
Comprehend It, Said Einstein When
His Daring Publishers Accepted It.

Asymmetric controversy (Collins and Pinch 1970s)

Claim of bias; rebutted by astronomers (RAS)

Einstein's universe

Einstein: apply GR to the Universe (1917)

Ultimate test for new theory of gravitation

Assumptions

Static universe (small velocities of the stars)

Mach's principle (metric tensor to vanish at infinity)

Isotropy and homogeneity (simplicity)

Boundary problem

Assume cosmos of closed curvature

But...no consistent solution

New term in field equations!

Cosmic constant - anti-gravity term

Radius and density defined by λ

$$G_{\mu\nu} = -\kappa T_{\mu\nu}$$

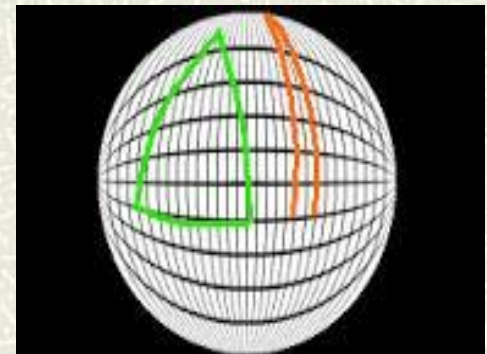
$$G_{\mu\nu} + \lambda g_{\mu\nu} = -\kappa T_{\mu\nu}$$

Doc. 43

Cosmological Considerations in the General Theory of Relativity

This translation by W. Perrett and G. B. Jeffery is reprinted from H. A. Lorentz et al., *The Principle of Relativity* (Dover, 1952), pp. 175–188.

It is well known that Poisson's equation $\nabla^2\phi = 4\pi K\rho$ (1) in combination with the equations of motion of a material point is not as yet a perfect substitute for Newton's theory of action at a distance. There is still to be taken into account the condition that at spatial infinity the potential ϕ tends



$$\lambda = \frac{\kappa\rho}{2} = \frac{1}{R^2}$$

Friedman's universe



Alexander Friedman
(1888 -1925)

Allow time-varying solutions (1922)

Assume homogeneity, isotropy, positive curvature

Two independent differential equations from GFE

$$G_{\mu\nu} + \lambda g_{\mu\nu} = -\kappa T_{\mu\nu}$$

$$\frac{3R'^2}{R^2} + \frac{3c^2}{R^2} - \lambda = \kappa c^2 \rho,$$

Evolving universes

Density of matter varies over time

$$\frac{R'^2}{R^2} + \frac{2RR''}{R^2} + \frac{c^2}{R^2} - \lambda = 0.$$

Negative spatial curvature (1924)

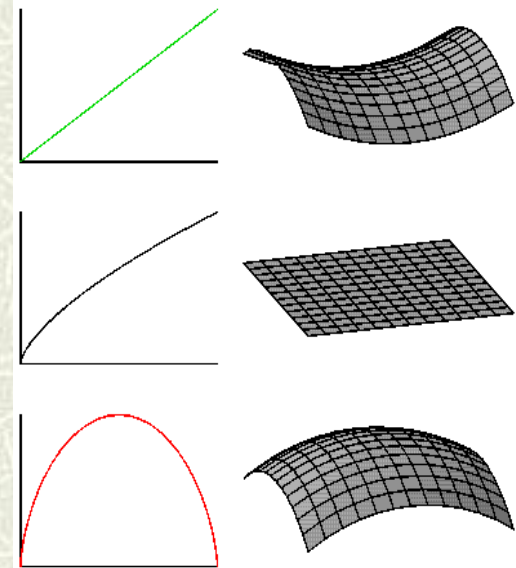
Cosmic evolution, geometry depends on matter

Overlooked by community

Considered 'suspicious' by Einstein

Mathematical correction, later retracted

~~"To this a physical reality can hardly be ascribed"~~



Lemaître's universe (1927)



✦ Expanding model of the cosmos from GR

Similar but not identical to Friedman 1922

$$3\frac{R'^2}{R^2} + \frac{3}{R^2} = \lambda + \kappa\rho$$

✦ Redshifts of galaxies = expansion of space?

Redshifts from Slipher, distances from Hubble

$$2\frac{R''}{R} + \frac{R'^2}{R^2} + \frac{1}{R^2} = \lambda - \kappa p$$

Fr Georges Lemaître

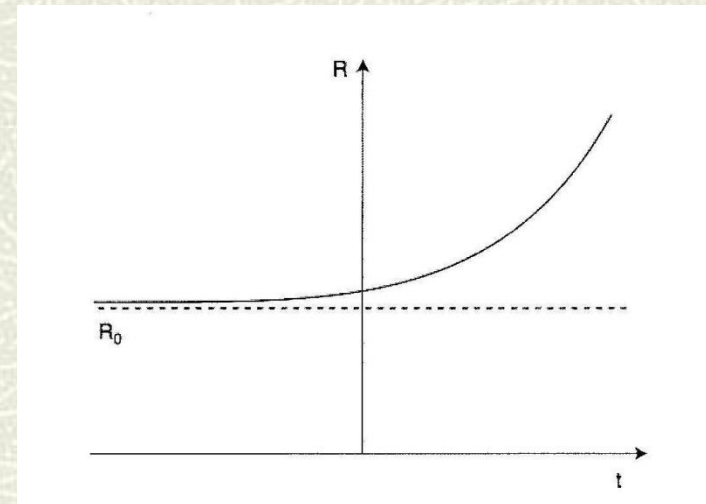
✦ Ignored by community

Belgian journal (in French)

Rejected by Einstein: "Vôtre physique est abominable"

✦ Lemaître's recollection (1958)

"Einstein not up-to-date with astronomy"



The watershed: Hubble's law (1929)



■ The redshifts of the spiral nebulae

Doppler shifts? (Slipher 1915, 1917)

■ The distances to the nebulae

Far beyond Milky Way (Hubble 1925)

■ A redshift/distance relation

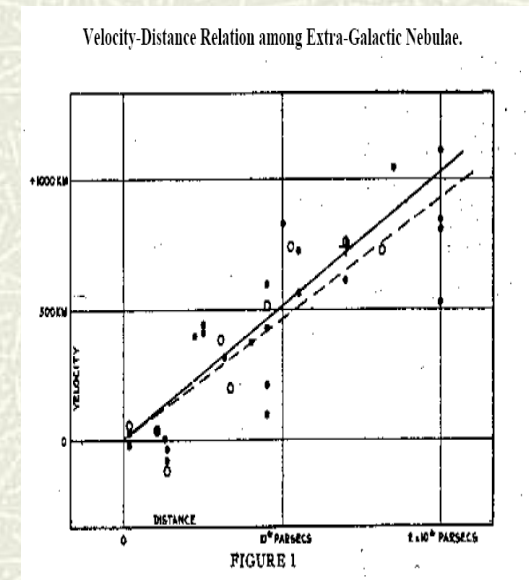
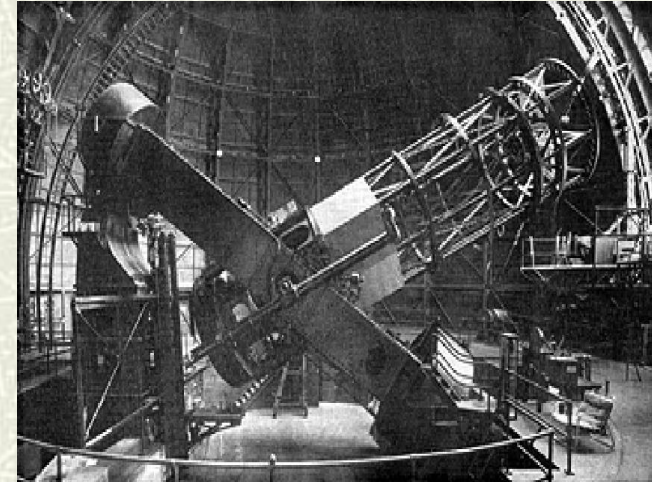
Linear relation (Hubble, 1929)

$$H = 500 \text{ km s}^{-1} \text{ Mpc}^{-1}$$

■ Landmark result in astronomy

Furthest galaxies receding the fastest

Link to relativity



The expanding universe (1930)

- **RAS meeting (May 1930)**

*If redshifts are velocities, and if effect is non-local
Hubble's law = expansion of space? (Edd., de Sitter)*

- **Dynamic model required**

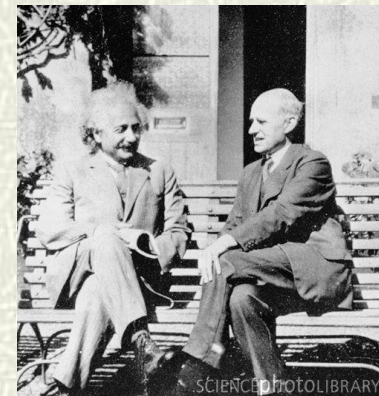
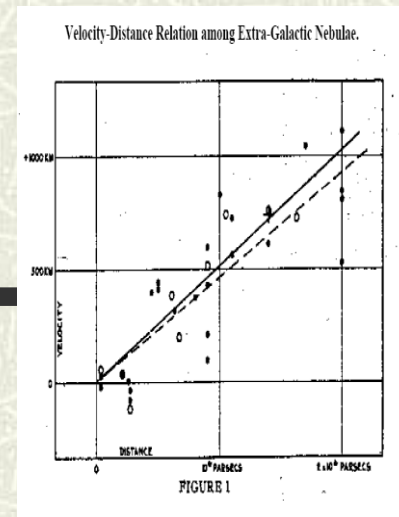
Static model unstable

- **Lemaître model adopted**

*1927 expanding model republished in English (1931)
Observational section omitted (rightly)*

- **Lemaître-Friedman cosmology accepted**

Time-varying radius, decreasing density of matter



Who discovered the expanding universe?

- # Einstein *Basic framework for theory*
- # Friedman *Expanding universe (theoretical)*
- # Hubble, Slipher *Recession of the galaxies (obs)*
- # Lemaitre *Observation + theory*



‘Hubble graph’ should be Hubble-Slipher graph

‘Hubble expansion’ should be Hubble-Lemaître expansion

Astronomers sceptical (Hubble)

Models of the expanding universe (1930 -)

- **Tolman (1930, 31)**

Expansion caused by annihilation of matter ?

- **Eddington (1930, 31)**

*On the instability of the Einstein universe
Expansion caused by condensation?*

- **de Sitter (1930, 31)**

Variety of expanding models

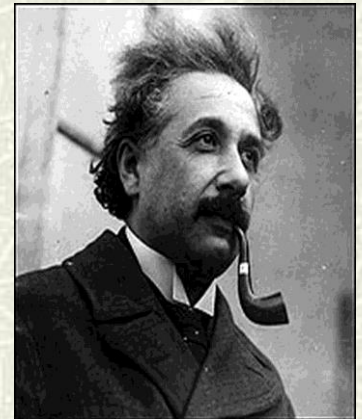
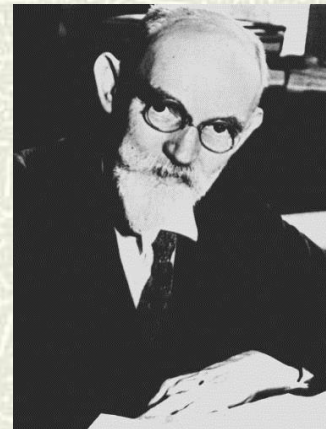
- **Heckmann (1931,32)**

Spatial curvature (not translated)

- **Einstein (1931, 32)**

Friedman-Einstein model $\lambda = 0, k = 1$

Einstein-de Sitter model $\lambda = 0, k = 0$



*If redshifts represent expansion...
If effect is global...*

A strange prediction: the big bang



Fr Georges Lemaître

- # **Lemaître 1931:** expanding U smaller in the past
- # Extrapolate to very early epochs
- # Extremely dense, extremely hot
- # Expanding and cooling ever since
- # ‘Fireworks beginning’ at $R = 0$?

Rejected by community (1930-60)

Simplified models

Timescale problem



Later called ‘The big bang’

A second line of evidence

Gamow and nuclear physics (1940s)

Student of Friedman

How were the chemical elements formed?

Problems with stellar nucleosynthesis

Elements formed in the infant hot universe?

Theory predicts $U = 75\%$ Hydrogen, 25% Helium

Agreement with observation

Victory for big bang model



Georges Gamow



Heavier atoms formed in stars

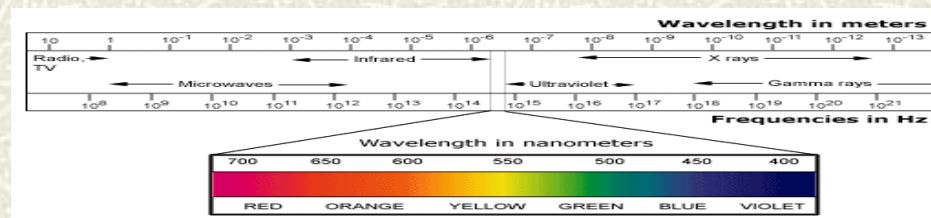
Bonus: cosmic background radiation

- Infant universe very hot
- Dominated by radiation
- Radiation still observable today?
Low temp, microwave frequency
- A fossil from the early universe!
Released when atoms formed (300,000 yr)

No-one looked



Alpher, Gamow and Herman



The steady-state universe (1948)

‡ Expanding but unchanging universe

Hoyle, Bondi and Gold (1948)

Disliked extrapolation to early epochs

Perfect cosmological principle?

‡ Requires continuous creation of matter

Very little matter required

No beginning, no age paradox

‡ Replace λ with creation term (Hoyle)

$$G_{\mu\nu} + C_{\mu\nu} = -k T_{\mu\nu}$$

‡ Improved version (1962)

$$G_{\mu\nu} + \lambda g_{\mu\nu} = k T (C_{\mu} + C_{\nu})$$



Bondi, Gold and Hoyle



Hoyle and Narlikar (1962)

Steady-state vs big bang (1950-70)

Optical astronomy (1950s)

Revised distances to the nebulae (Baade, Sandage)

New timescale of expansion

Radio-astronomy (1960s)

Galaxy distributions at different epochs

Cambridge 3C Survey (Ryle)

Nucleosynthesis of light elements

Gamow et al. 1948

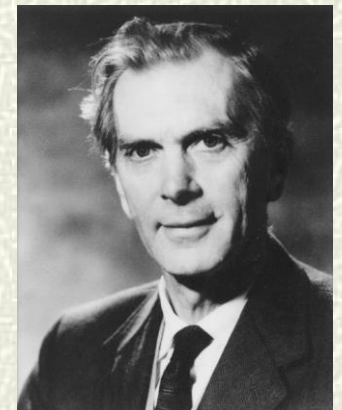
Cosmic microwave background (1965)

Low temperature, low frequency

Remnant of young, hot universe



Allen Sandage



Martin Ryle

Cosmic background radiation (exp)

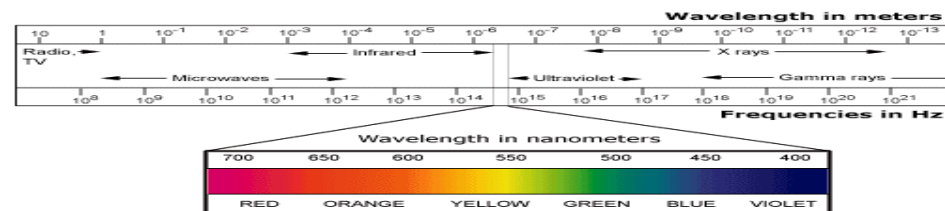
CMB discovered accidentally

- # Universal signal (1965)
- # Low frequency (microwave)
- # Low temperature (3K)



Penzias and Wilson (1965)

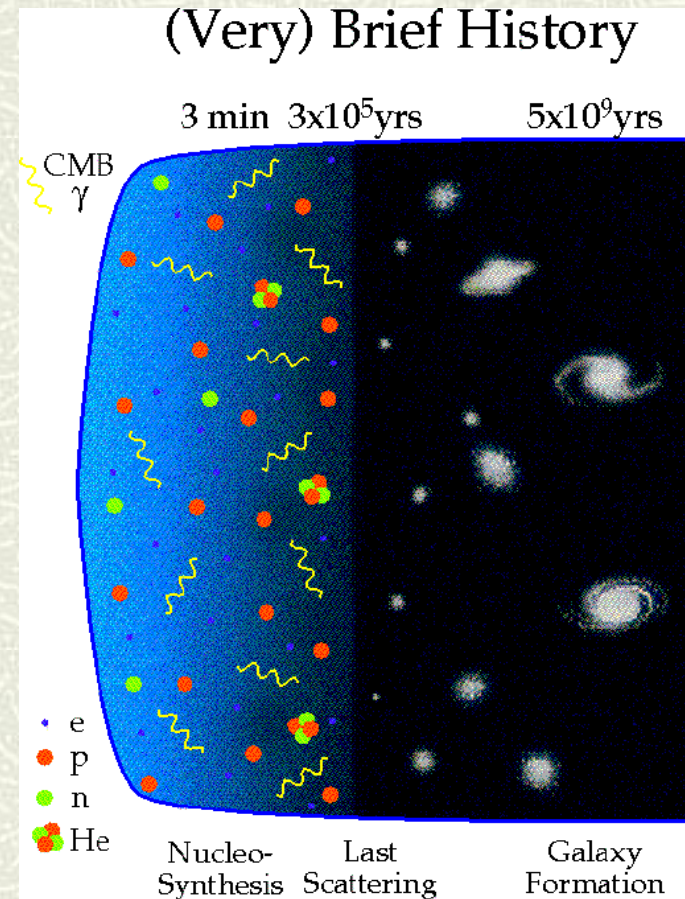
*Echo of **Big Bang!***



The big bang – evidence

1. The expansion of the U
2. The abundance of H and He
3. The distribution of the galaxies
4. The cosmic microwave background

How did it start?



Cosmology in the 1970s, 1980s

Astronomical search for cosmic parameters

Density of matter, rate of expansion

The search for two numbers

H_0 , Ω_M

q_0

$$q_0 = \frac{\Omega_M}{2}$$

Hypothesis of dark matter

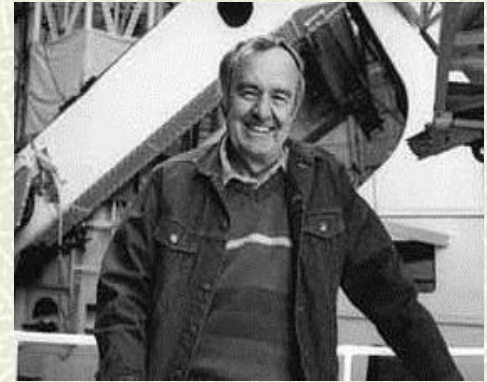
Dynamics of galaxy rotation

Some matter seen only by its gravitational effect

Problems

Density of matter too small to close universe

But no evidence of spatial curvature!



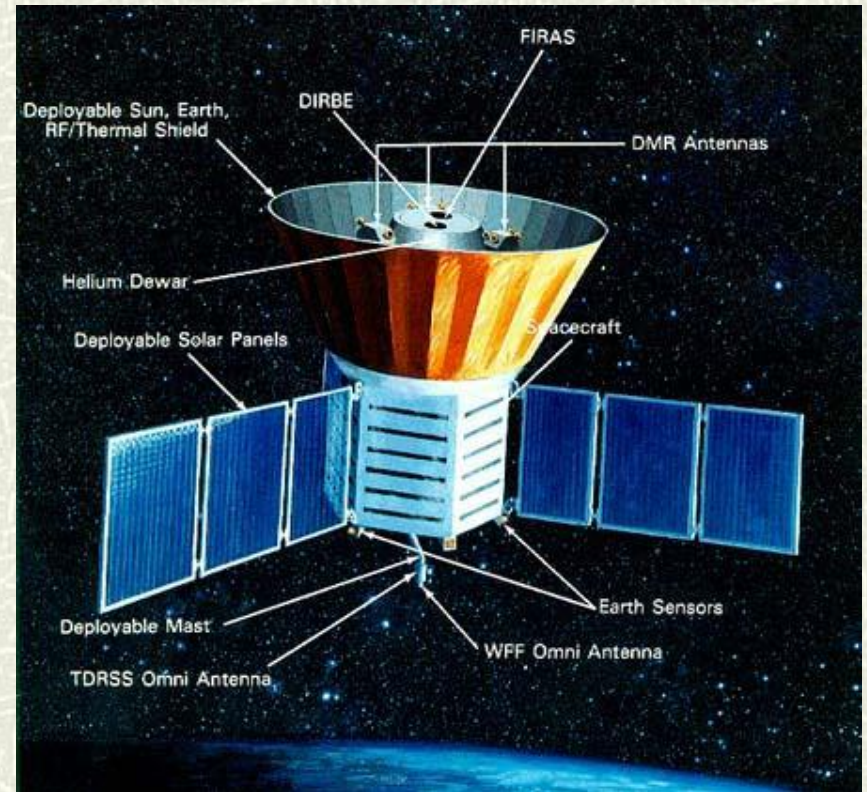
Allan Sandage



Vera Rubin

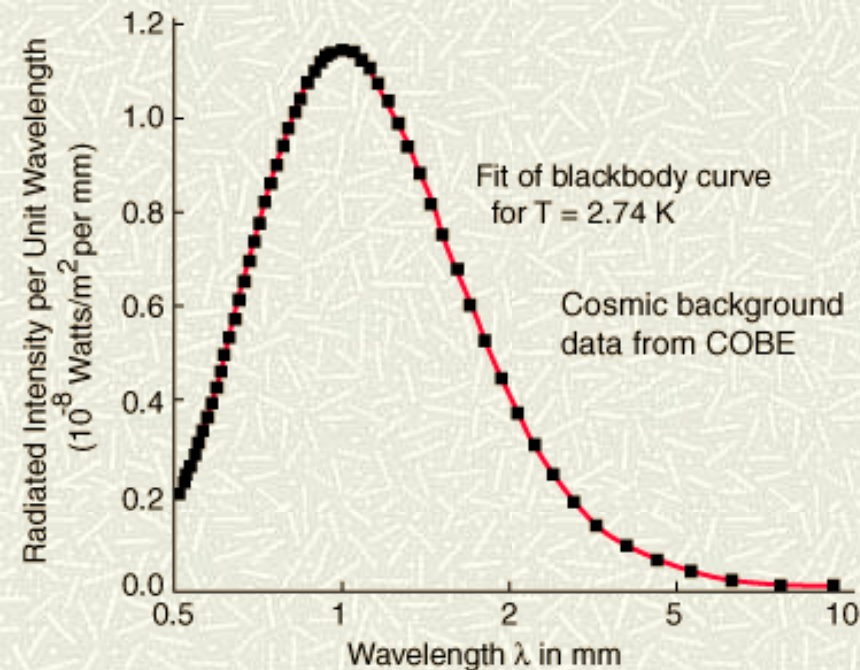
Modern measurements of the CMB

- Details of background radiation
 - Full spectrum
 - Comparison with theory
 - Perturbations?
-
- *Ground telescopes*
 - *Balloon experiments*
 - *Satellite experiments*



COBE satellite (1992)

Cosmic background radiation



COBE (1992)

Nobel Prize 2006

- **Expected temperature**
- **Expected frequency**
- **Perfect blackbody spectrum**
- *radiation too uniform?*
- *perturbations < 1 in 10^5 ?*
- *galaxy formation?*

Problems



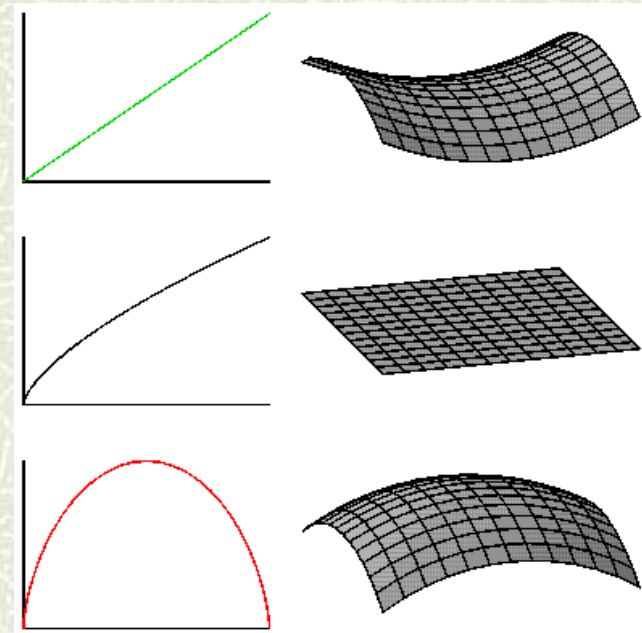
Background radiation raised new questions

Robert Dicke

- ✦ Horizon problem *why so uniform?*
- ✦ Galaxy problem *how did galaxies form?*
- ✦ Flatness problem *fine balance?*

Conflict between theory and experiment

Astrophysics: $\Omega = 0.3$

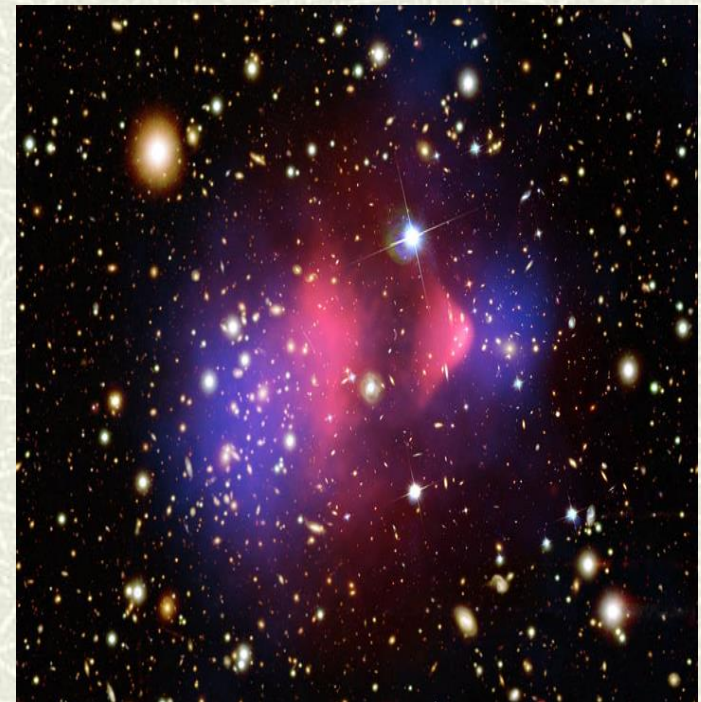


Dark Matter

- # First suggested in 1930s
- # Stellar motion

*normal gravitational effect but
cannot be seen directly*

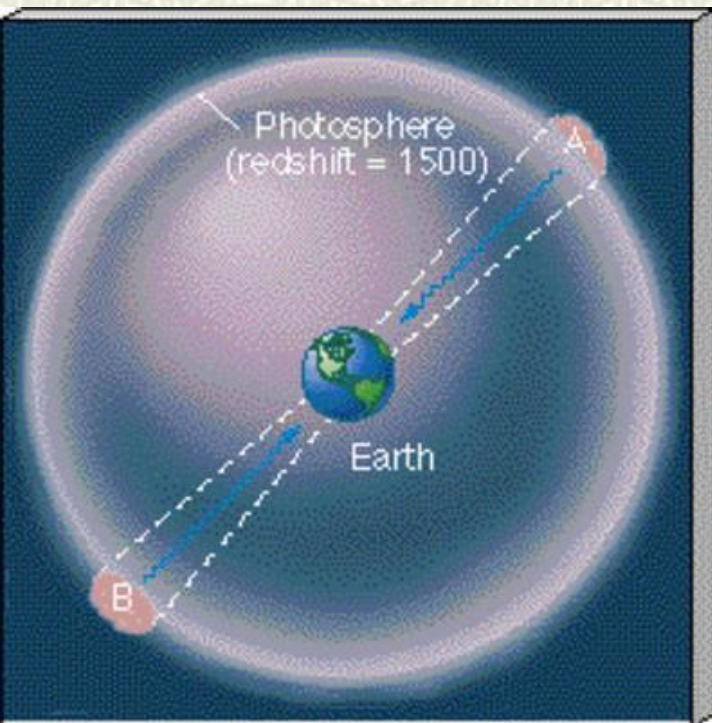
- # Explains motion of stars
- # Explains motion of galaxies
- # Explains gravitational lensing



Matter = OM (20%) + DM (80%)

$\Omega = 0.3 ?$

The horizon problem



- Two distant regions of microwave background have similar temps

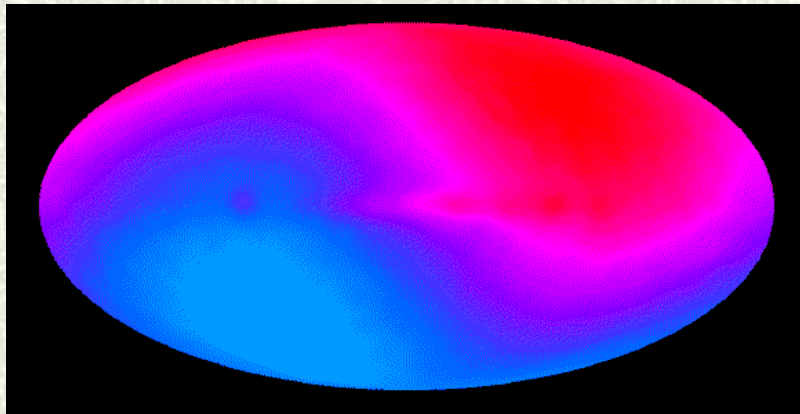
Why?

Too far apart to be causally connected

- Finite speed of light
- Finite age of cosmos

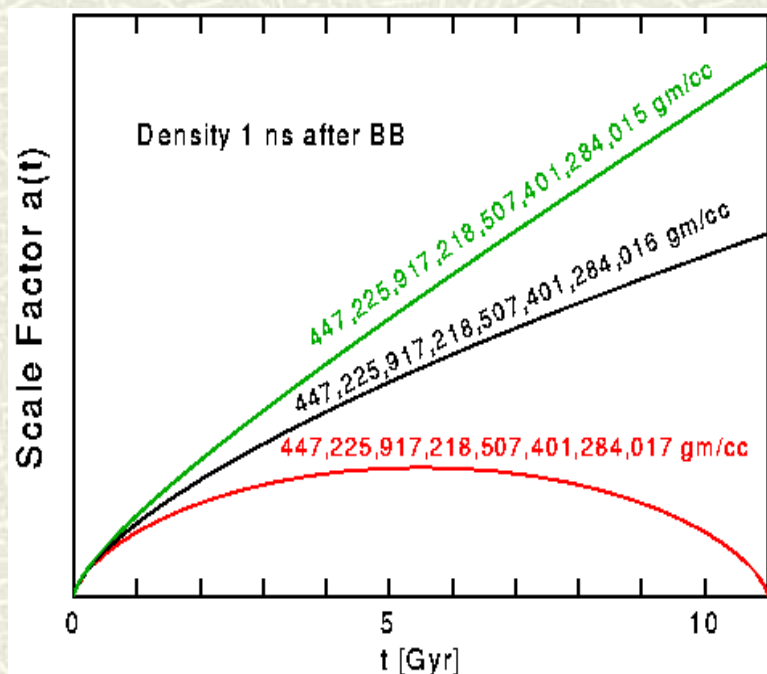
Something wrong with lookback

Galaxy formation problem



- ✦ Microwave background smooth on large scale
- ✦ No deviations from homogeneity obvious
(1 in 10,000)
- ✦ How did slight perturbations become galaxies?

The flatness problem



At $t = 1 \text{ s}$, $\Omega = 1$ to within $1:10^{15}$)

Slightest deviation from flatness



runaway expansion or crunch

Not observed

Why so finely balanced initially?

$$\Omega = 1?$$

Astrophysics: $\Omega = 0.3$?

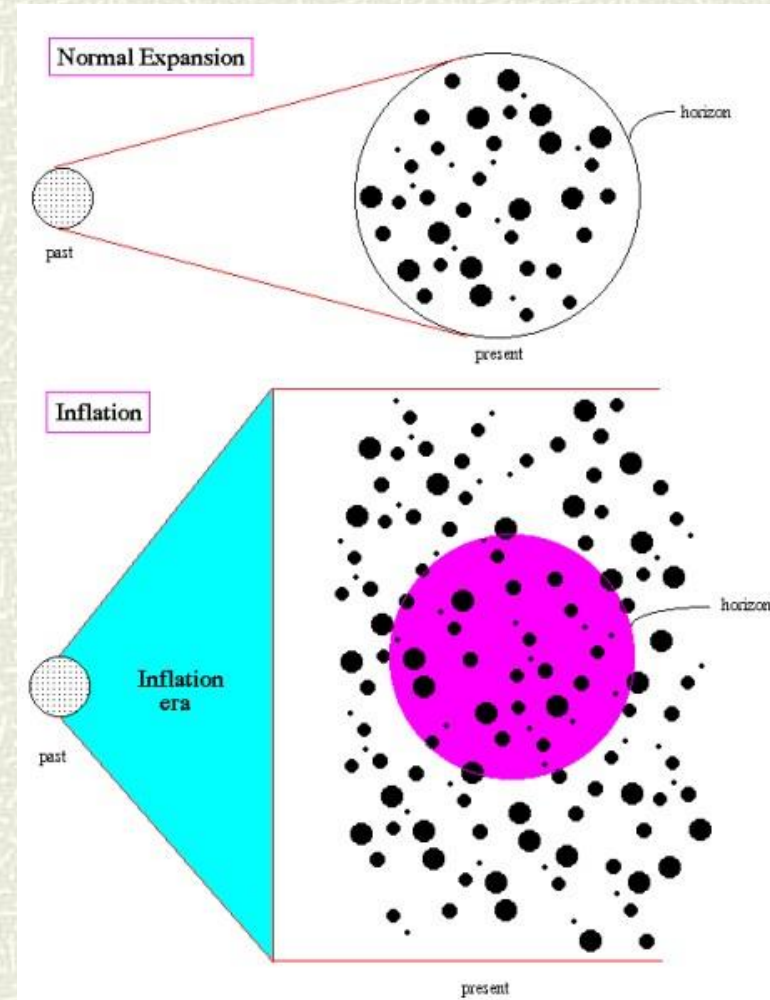
The Theory of Inflation (1981)

- Initial **exponential expansion**
- Driven by *phase transition*

Repulsive force

- Expansion of 10^{35} in 10^{-32} s
- Smooths out inhomogeneities
- Smooths out curvature

'No hair' universe

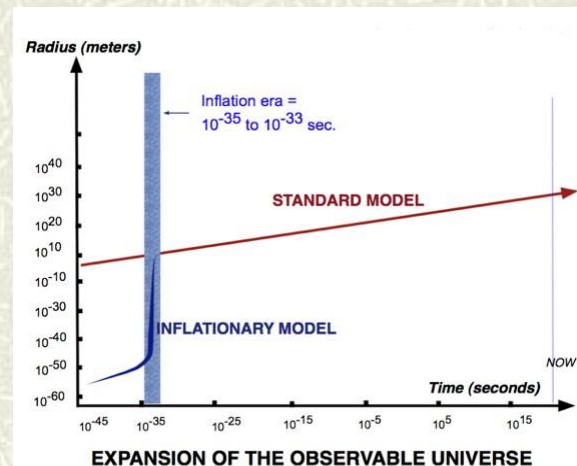
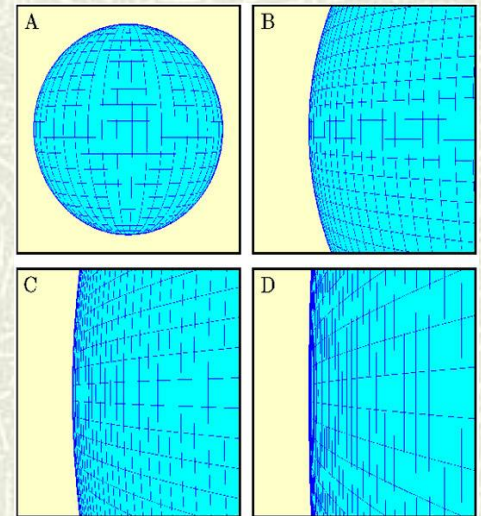


The inflationary universe

- ⌘ Solves horizon problem
Early U incredibly small
- ⌘ Solves flatness problem
Geometry driven towards flatness
- ⌘ Mechanism for galaxy formation
Quantum fluctuations inflated

$$\Omega = 1 ?$$

Conflict between theorists and experimentalists



Problems with standard model (1980-1998)

✦ Einstein-de Sitter: $\lambda = 0, k = 0$

✦ Flatness prediction (Dicke, inflation)

$$k = 0$$

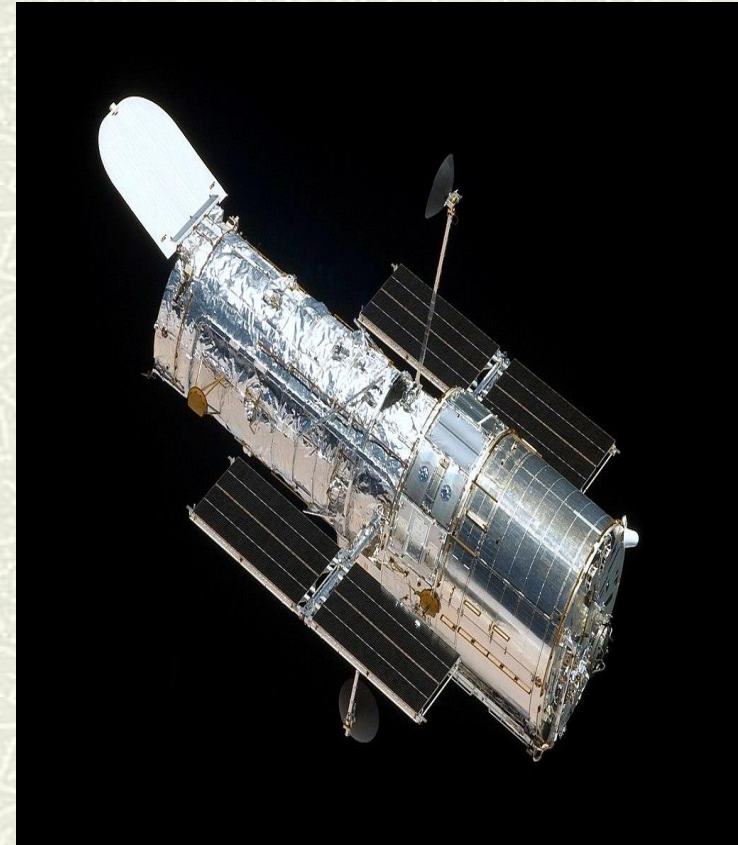
✦ But: $\Omega_M = 0.3$

✦ But: new timescale problem from HST

$$H = 87 \text{ kms}^{-1} \text{ Mpc}^{-1} : t = 11.2 \times 10^9 \text{ yr}$$

too young

Is model wrong? Is $\lambda \neq 0$?



*Hubble Space Telescope
1990*

The accelerating universe (1998)

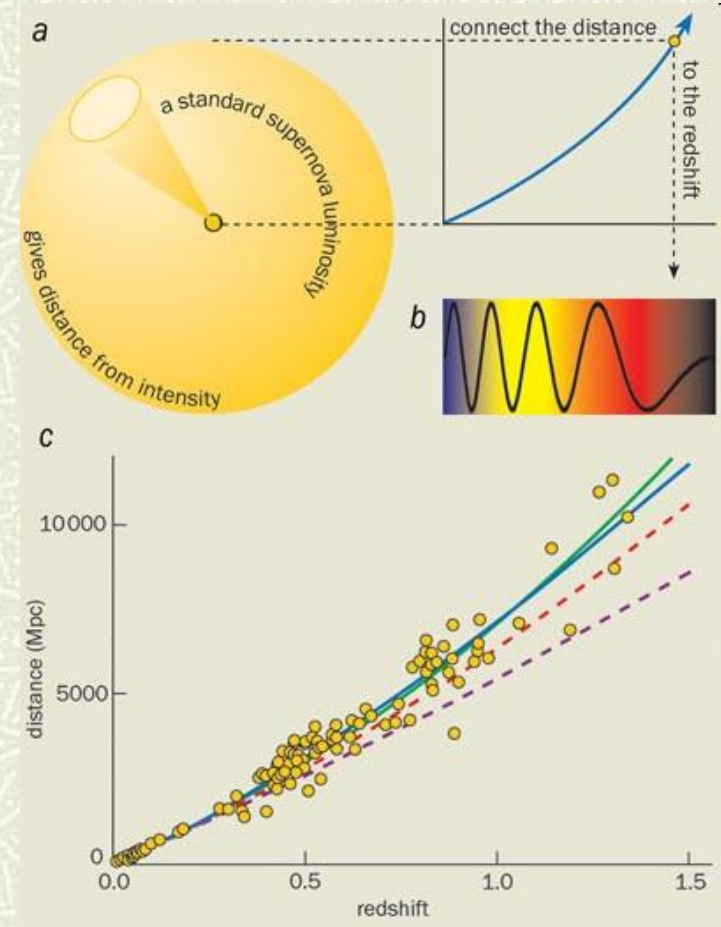
- Supernovae as standard candles (1998)

- Furthest galaxies too far away

- Expansion accelerating

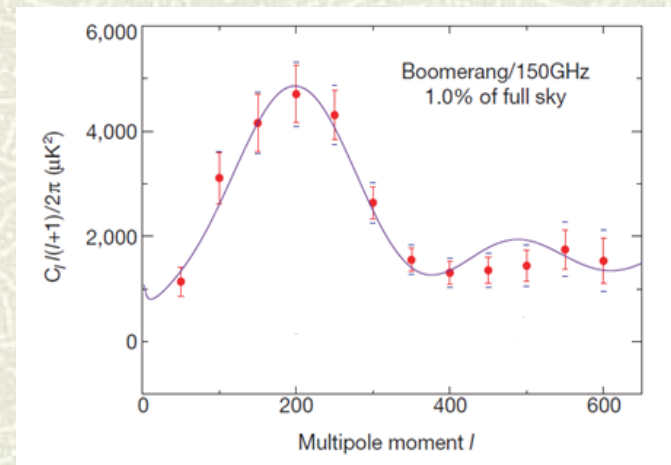
- Implies repulsive energy
Dark Energy ($\lambda \neq 0$)

- Possible support for inflation ($\Omega = 1$?)



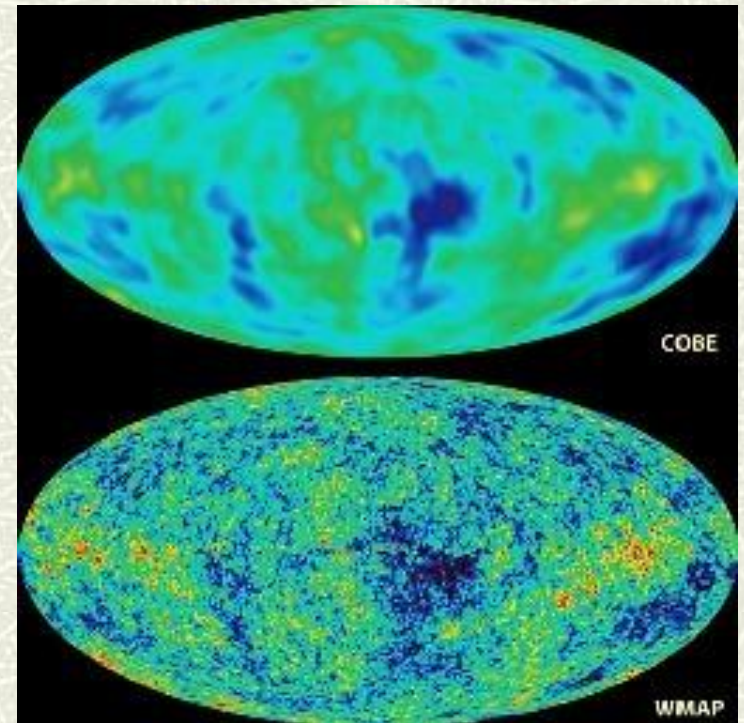
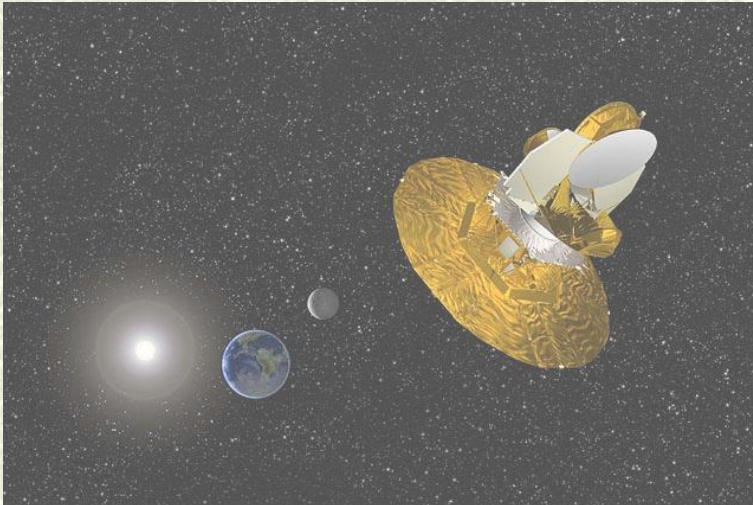
The flat universe (2000)

- Balloon measurements of CMB
- The BOOMERANG experiment
- The MAXIMA experiment
- Geometry is flat ($\Omega = 1$)
Implies dark energy component
$$\Omega_M + \Omega_\lambda = 1 \quad (\Omega_\lambda = 0.7)$$
- Support for inflation



WMAP Satellite (2002)

- Details of *CMB* spectrum
- Details of galaxy formation
- Details of flatness of *U*
- Details of dark energy



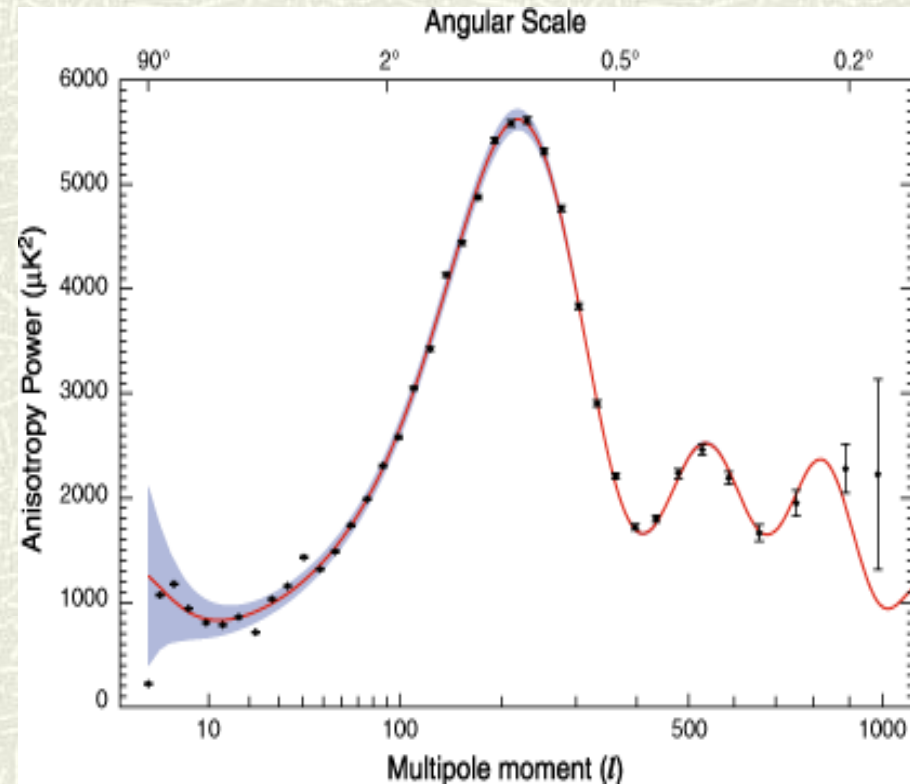
Cosmic microwave background

WMAP measurements of CMB (2005)

- Spectrum of T variations
- Geometry is flat (to 1%)
- Dark energy 74%

Strong support for dark energy

Strong support for inflation

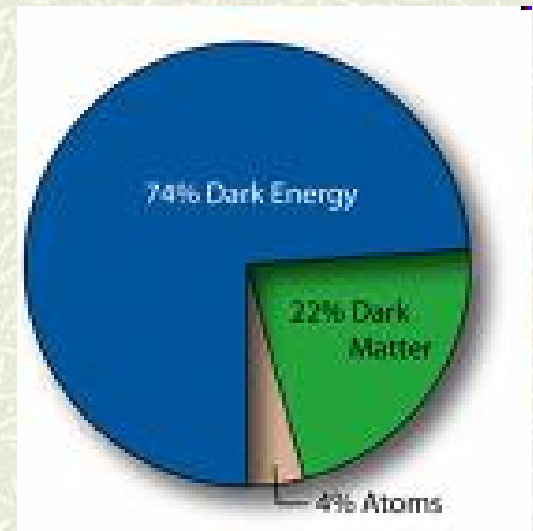


Fit to theory

Modern big bang model: Λ -CDM

A flat, accelerating universe containing matter, dark matter and dark energy

1. Ordinary matter: 4% (astrophysics)
2. Dark matter: 22% (astrophysics)
3. Dark energy : 74% (supernova, CMB)

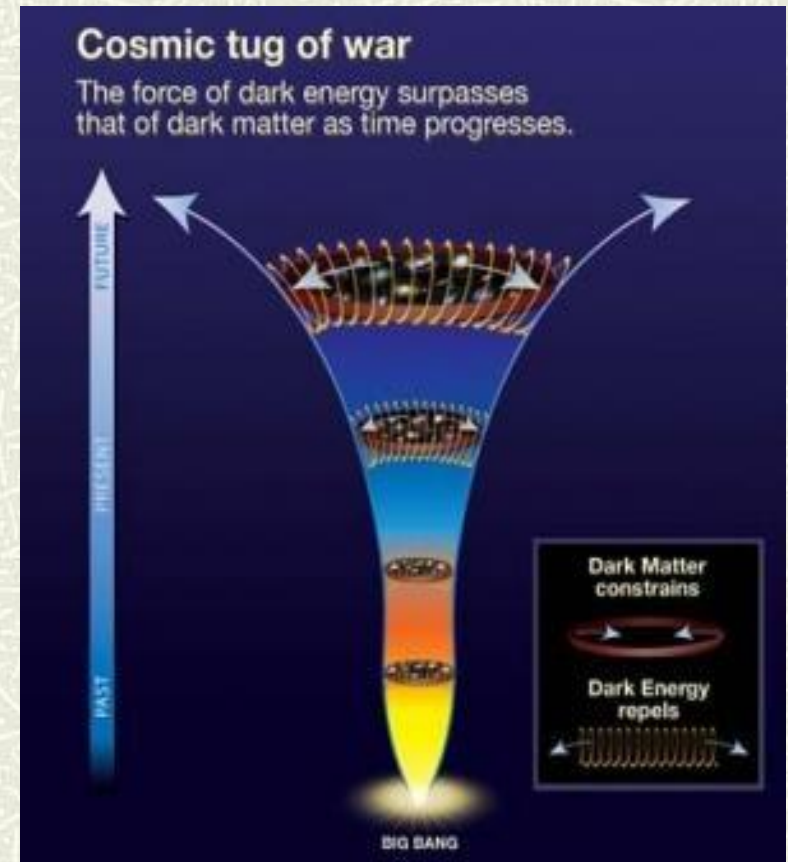


Λ CDM

$$\Omega = 1$$

Cause of acceleration: dark energy

- Predicted by relativity
- Cosmological constant
- Natural tendency of space to expand
- Energy of vacuum?
- Why so small?
- Why of similar density to matter?
- Not well understood
- Fate of universe?



New results: Planck Satellite (ESA, 2013)

1. Improved sensitivity

$$\Delta T/T \approx 1 \times 10^{-6}$$

2. Full spectrum of T anisotropy

New acoustic peaks :scale invariance?

Accurate values for Ω_{Λ} , Ω_{M}

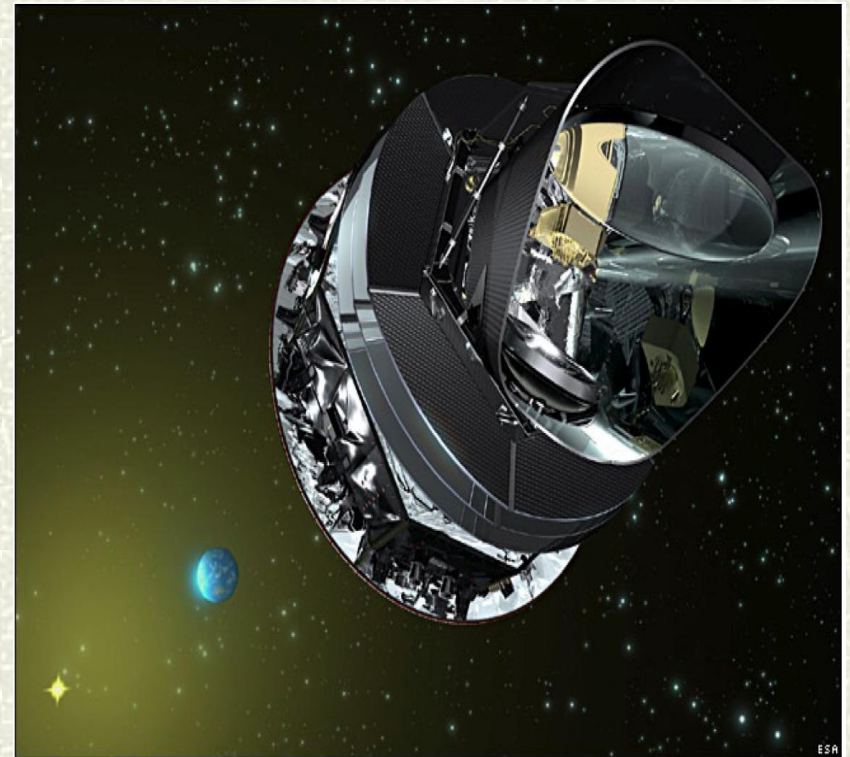
3. Gravitational lensing

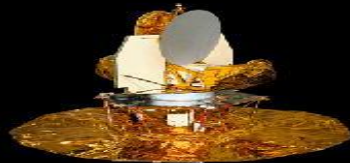
Remove degeneracies

4. Polarization measurements

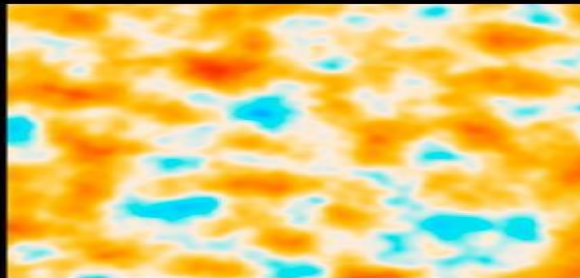
E-modes: fluctuations

B-modes: gravity waves?

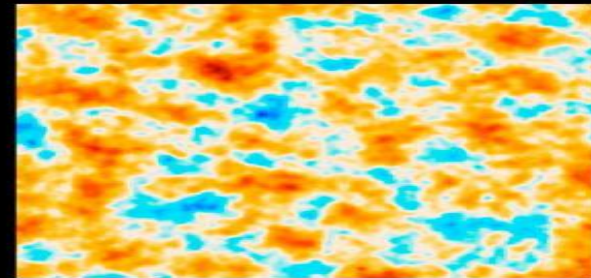




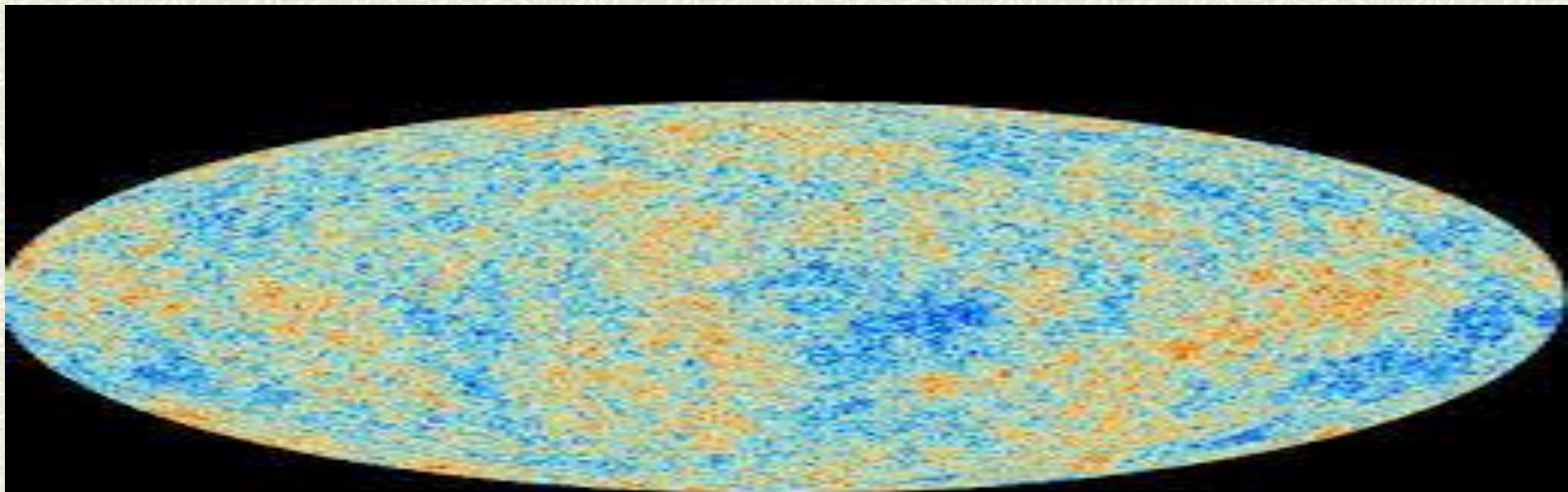
COBE



WMAP



Planck

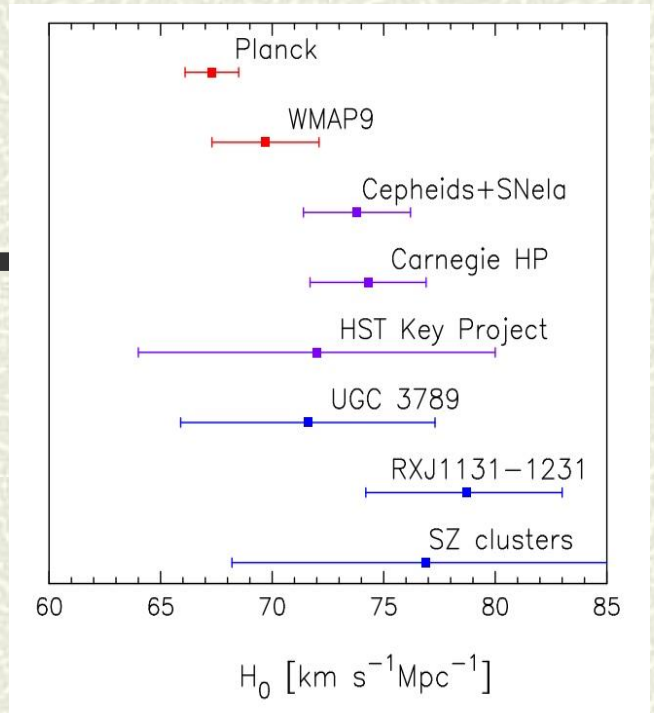


Planck results (2013)

1. New Hubble constant

$67.3 \pm 1.2 \text{ km/s/MPC}$

$\text{Age} = 13.8 \text{ billion yr}$

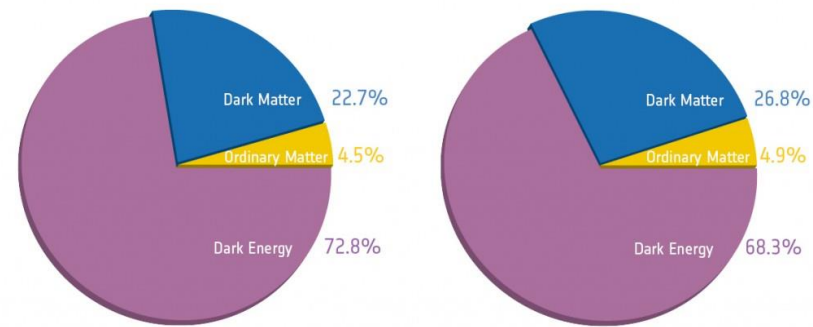


2. Curvature ; flat

$$\Omega_k = -0.0005 \pm .07$$

3. New mass/energy parameters

$$\Omega_\Lambda = 68, \Omega_{\text{DM}} = 27, \Omega_{\text{OM}} = 4.9 \%$$



Before Planck

After Planck

The big bang - problems

- ⌘ Nature of dark energy?

Role in BB?

- ⌘ Which model of inflation?

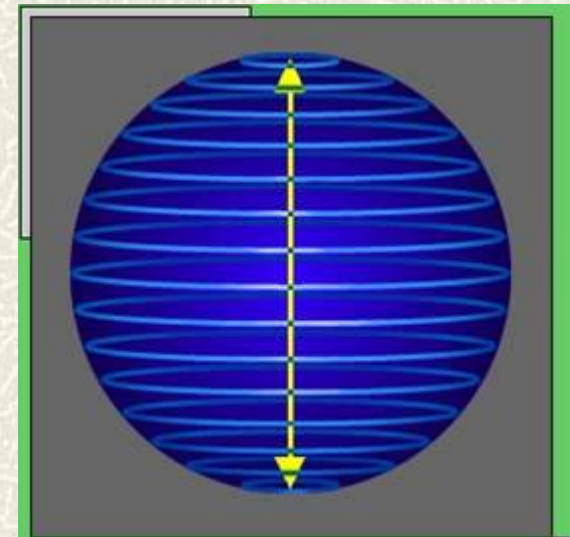
The multiverse?

- ⌘ The singularity problem

What banged?

What does time zero mean?

No-boundary universe



The singularity: a cyclic universe?

- ⌘ Breakdown at time zero
- ⌘ No model of bang itself
- ⌘ Multiple bangs?
- ⌘ Colliding branes
- ⌘ Prediction of string theory
- ⌘ Cyclic universe
- ⌘ Eternal universe



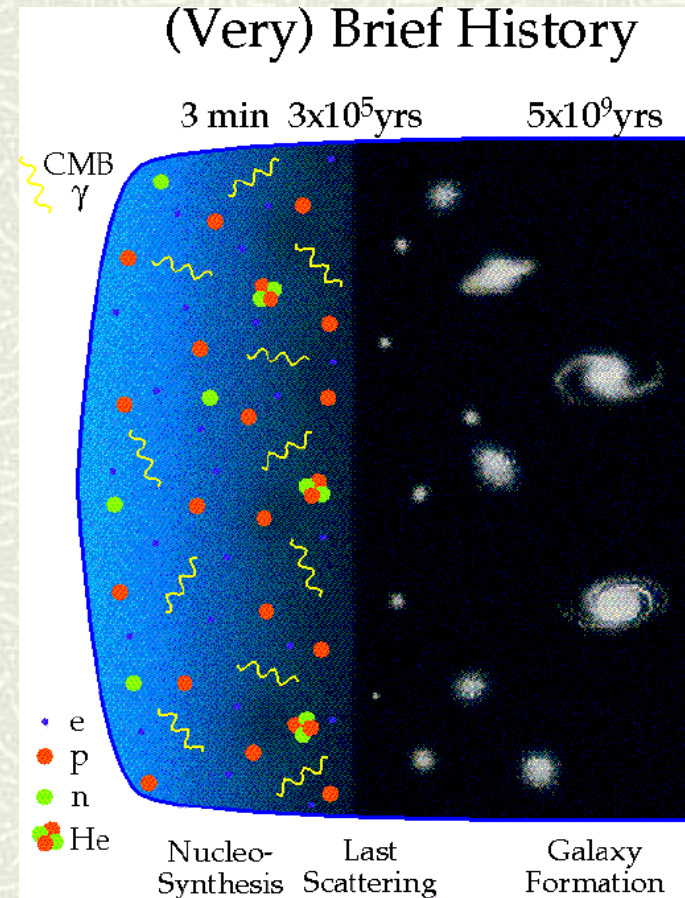
Cyclic universe

Tests? Non-Gaussianities in CMB

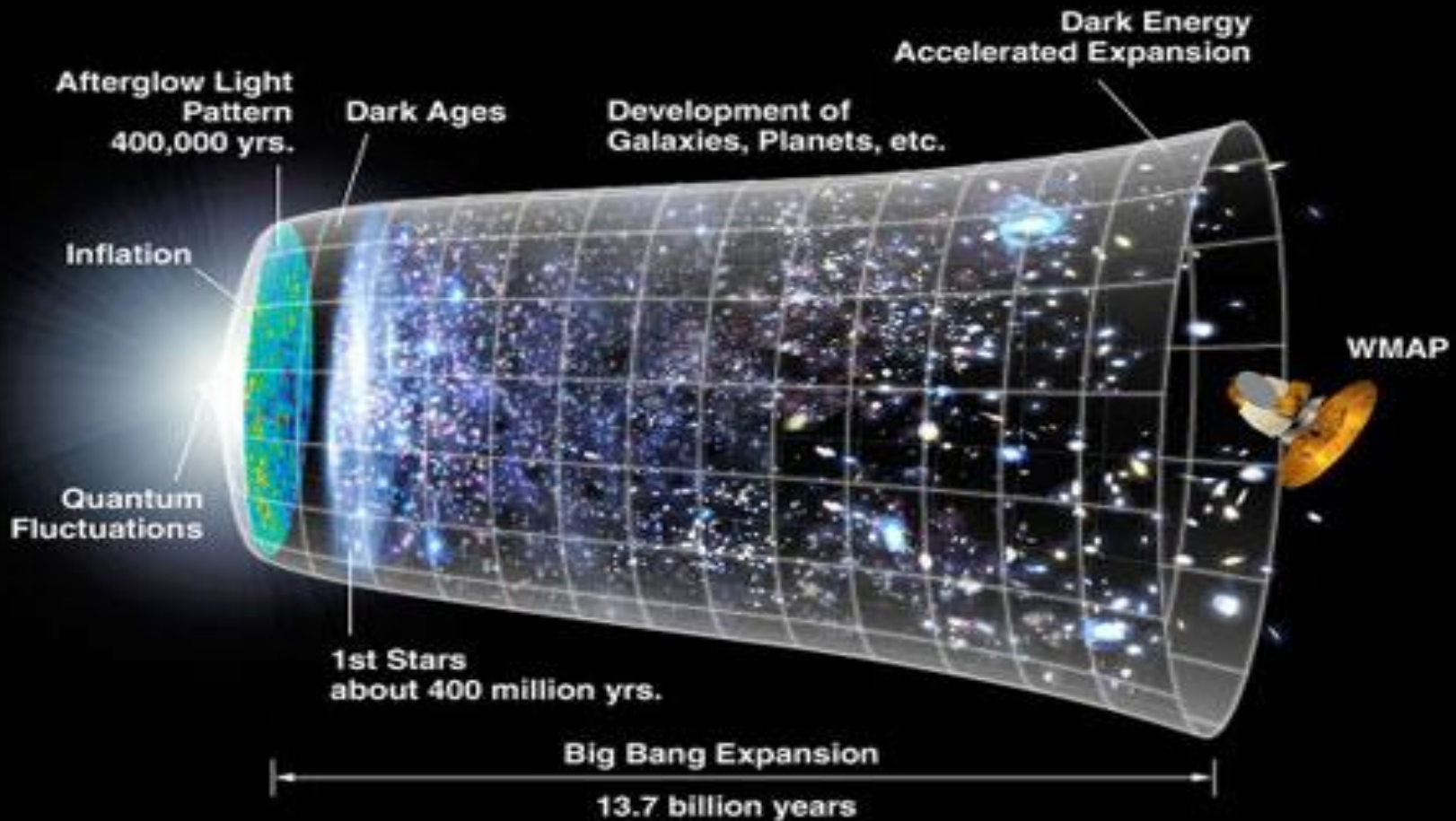
The big bang – evidence

1. The expansion of the U
2. The abundance of H and He
3. The distribution of the galaxies
4. The cosmic microwave background

How did it start?



The big bang model



Today's cosmological puzzles

Characteristics of background radiation

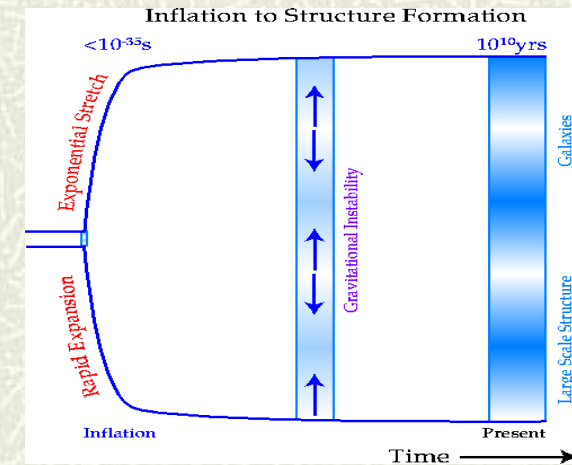
Isotropy, homogeneity, flatness (1970-80)

The theory of inflation (1981)

Exponential expansion within first second?

Explanation for homogeneity, flatness, galaxy formation

Which model of inflation? Bubble inflation?



Dark energy (1998)

Observation of accelerated expansion

The return of the cosmological constant

U mainly composed of dark energy

$$G_{\mu\nu} + \lambda g_{\mu\nu} = -\kappa T_{\mu\nu}$$

Nature of DE unknown

