Redshifts and the expanding universe

Paradigm shift or slow dawning?

Cormac O'Raifeartaigh FRAS

A drama in three acts?

★ A brief history of observation (1912-1931)

The redshifts of the spiral nebulae (Slipher)
The distances to the nebulae (Hubble)
The Hubble graph of 1929

★ A brief history of theory (1915-1931)

The 'static' universes of Einstein and de Sitter
The dynamic universes of Friedman and Lemaître

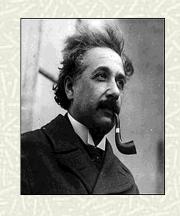
An expanding universe? (1930) *Explorations of a dynamic universe (1930-35)*

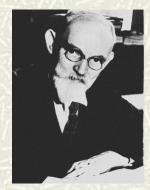
Paradigm shift or slow dawning?

■ New findings *Einstein's steady-state model (1931)*













The starry nebulae

- **** Observed by Marius (1614), Halley, Messier**Cloudy structures; not planets or stars
- **Island universes? Kant, Laplace (1755-96)**Galaxies of stars at immense distance?

 Are stars born in the nebulae?
- ₩ Wilhem Herschel36-inch reflecting telescopeCatalogue of a thousand (1786)
- **□ Earl of Rosse**72-inch reflecting telescope (1845)Some nebulae have spiral structure, stars

Problem of resolution, distance







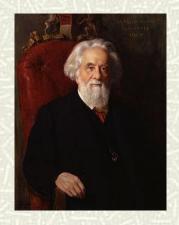
The spectra of the nebulae

- **♯** Photography and spectroscopy (19th cent)

 Emission and absorption lines of celestial objects
- **Composition of the stars**William Huggins
- **Motion of the stars: Doppler effect**William Campbell
- **♯** Spectroscopy of spiral nebulae?

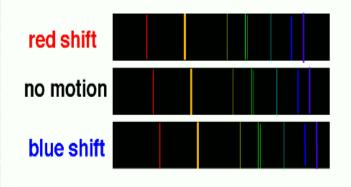
 Composition of nebulae?

 Motion of nebulae?
- **■** Difficult to resolve





$$\Delta \lambda / \lambda = v/c$$



Slipher and the nebulae

- Analyse light of the spiral nebulae? (1909)

 Lowell Observatory; evolving solar system?
- Slipher reluctant
 24-inch refractor: larger telescopes failed
- Experiments with spectrograph camera

 Good results with fast camera lens
- Clear spectrum for Andromeda nebula (1912) Significantly blue-shifted; approaching at 300 km/s?
- Many spiral nebulae red-shifted (1915)

 Standing ovation (AAS, 1914)

 Attended by Hubble



V.M. Slipher



Redshifts of the nebulae

- Spectra of 25 spirals (1917)
 Large outward velocities
 Some receding at 1000 km/s
- Much faster than stars
 Gravitationally bound by MW?
- Island universe debate

 "Island universe hypothesis gains favour"
- Faintest spectra most redshifted

 Evidence of expansion? (retrospective)
- 41 redshifts by 1922

 Published by Eddington, Strömberg
 What did they mean?



 $\Delta \lambda / \lambda = v/c$

RADIAL VELOCITIES OF TWENTY-FIVE SPIRAL NEBULE.

Nebula,	Vel.	Nebula.	Vel.
N.G.C. 221	- 300 km.	N.G.C. 4526	+ 580 km
224	- 300	4565	+1100
598	- 260	4594	+1100
1023	+ 300	4649	+1090
1068	+1100	4736	+ 290
2683	+ 400	4826	+ 150
303 I	- 30	5005	+ 900
3115	+ 600	5055	+ 450
3379	+ 780	5194	+ 270
3521	+ 730	5236	+ 500
3623	+ 800	5866	+ 650
3627	+ 650	7331	+ 500
4258	+ 500		888 825

Act II: General relativity

Space+time = space-time

Spacetime dynamic (1905)

■ Spacetime distorted by mass

Distortion causes other mass to move (1915)



$$G_{\mu\nu} = \frac{8\pi G}{c^4} T_{\mu\nu}$$

■ Dyson/Eddington expeditions (1919)

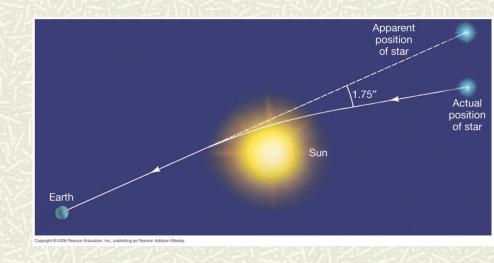
Measure bending of light?

Successful result

General relativity well-known



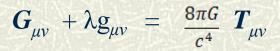
Albert Einstein



Relativity and the cosmos

Einstein model (1917)

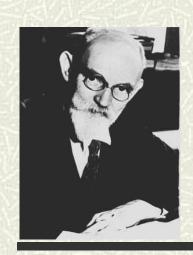
- # Homogenous fluid of uniform density
- # Equations predict dynamic universe
- No evidence for such a universe
- **♯** Add cosmic constant 'static'
- # Closed curvature, finite radius



De Sitter (1917)

- # 'Empty' universe
- **♯** Apparently static (co-ordinate system)
- **#** Cosmic constant determined by curvature of space
- **■** Redshifts due to time dilation/matter





Disliked by Einstein: Mach's principle

Redshifts and the de Sitter universe

Karl Wirtz (1922,24)

Redshifts of nebulae increasing with distance Dispersal effect? $v = 2200 - 1200 \log (Dm)$

Ludwik Silberstein (1924) ★

Relation between redshifts, distance, curvature $\Delta \lambda / \lambda = +/- r/R$ (global clusters)

Knut Lundmark (1924,25)

Velocity against distance; clusters, nebulae

Gustav Strömberg (1925)

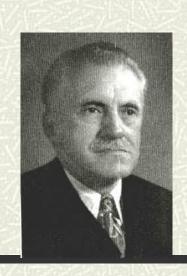
Vel/dist relation for globular clusters, nebulae?











Friedmann models

Allow <u>time-varying solutions</u> to the field equations

Expanding, contracting universes

Allow cosmic constant



Alexander Friedmann 1888 -1925

Geometry, evolution depends on matter

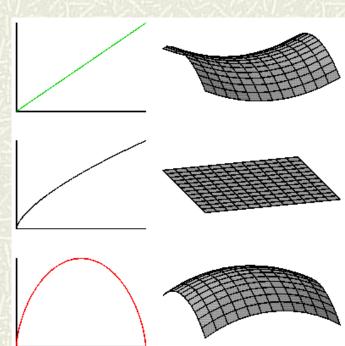
Positive curvature (1922)

Hyperbolic curvature (1924)

To be decided by astronomy

♯ Disliked by Einstein

Correction and retraction



Ignored by community

The distances of the nebulae (1925)

♯ Hooker telescope (Mt Wilson)

100-inch reflector (1917)

Edwin Hubble (1921)

Ambitious and dedicated astronomer

■ Resolved Cepheid stars in nebulae (1925)

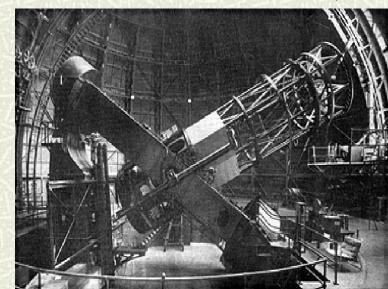
Leavitt's period-luminosity relation
Standard candle

♯ Spirals beyond Milky Way

Beginning of end of 'Great Debate'
Nebulae = galaxies



Edwin Hubble (1889-1953)



A redshift/distance relation (1929)

Is there a redshift/distance relation for galaxies?

<u>Motivation:</u> establishing distance to the galaxies

Combine 24 nebular distances with redshifts

Redshifts from Slipher: not acknowleged

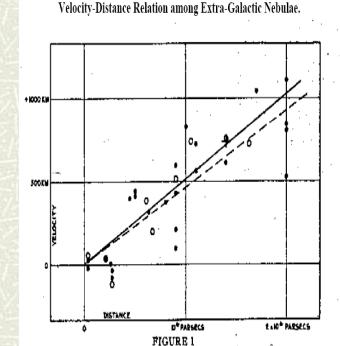
Approx linear relation (Hubble, 1929)

Some errors (Peacock)

Most important point not shown

What do the redshifts mean?

Reference to de Sitter universe



 $H = 585 \text{ kms}^{-1} \text{Mpc}^{-1}$

Act III An expanding universe? (1930-)

• RAS meeting (1930)

Eddington, de Sitter
Redshift/distance relation of the nebulae
Static models don't fit
New model required

Letter from Lemaître

Reminds Eddington of his <u>1927 model</u> Eddington, de Sitter impressed

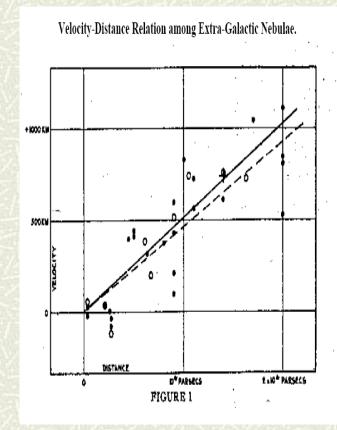
Expansion of space?

Considered by many theoreticians

If redshifts are velocities (Zwicky)

If effect is non-local

Not accepted by astronomers (Hubble)



Expansion of space?

Lemaître's universe (1927,31)



Rate of expansion from ave. distance and redshift H = 585 km/s/Mpc



de Sitter model not static (1925)

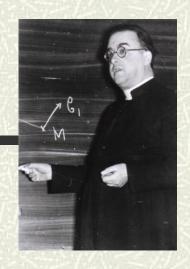
New evolving solution: Einstein \rightarrow de Sitter

■ No beginning: indefinite age

Starts from Einstein universe at $t = -\infty$

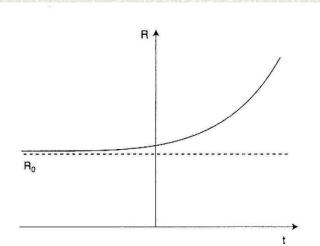
♯ Rejected by Einstein

An idea whose time had not yet come



Fr Georges Lemaître

Not an empirical law Edited in 1931 translation



Dynamic cosmic models (1930-32)

• Eddington (1930, 31)

On the instability of the Einstein universe The Eddington-Lemaître model Expansion caused by condensation?

• de Sitter (1930, 31)

Further remarks on the expanding universe Expanding universes of every flavour

• Tolman (1930, 31)

On the behaviour of non-static models Expansion caused by annihilation of matter?

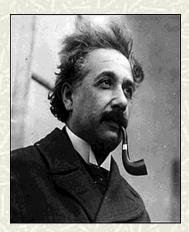
• Einstein (1931, 32)

Friedmann-Einstein model $\lambda = 0$, k = 1Einstein-deSitter model $\lambda = 0$, k = 0









If redshifts represent velocities...
If effect is non-local

Tolman's 'annihilation' model

♯ Non-static line element (1930)

Einstein, de Sitter models ruled out

$$ds^{2} = -\frac{e^{2kt}}{\left(1 + \frac{4r^{2}}{R^{2}}\right)^{2}} \left(dx_{1}^{2} + dx_{2}^{2} + dx_{3}^{2}\right) + c^{2}dt^{2}$$

Cause of cosmic expansion?

General evolutionary process

Transformation of matter into radiation

Rate of transformation

From Hubble's law and from stellar physics

May have influenced Einstein

Steady-state model

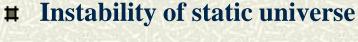


$$\frac{\delta\lambda}{\lambda} = k\Delta l$$

$$\frac{1}{M}\frac{dM}{dt} = -3k$$

$$k = 5 \times 10^{-10} \text{ yr}^{-1}$$

Einstein's 1931 model (F-E)



Eddington's paper



Expanding cosmos

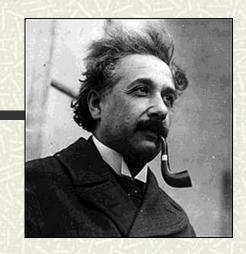
Remove cosmic constant?



Time-varying universe, k = 1, $\lambda = 0$

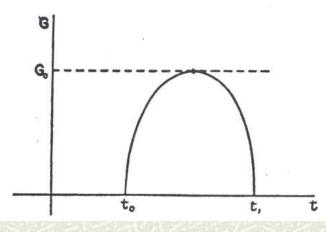
Age and singularity problems

Attributes to limitations of theory



Friedmann-Einstein universe Transl. O'Raifearaigh et al

$$\left(\frac{dP}{dt}\right)^2 = c^2 \frac{P_0 - P}{P}$$



Einstein's 1931 model (F-E)



Use Hubble parameter $P \sim 10^8 \ light-years, \ \rho \sim 10^{-26} \ g/cm^3$



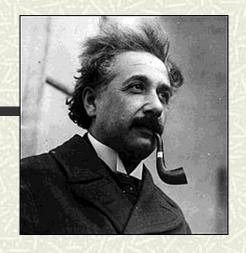
 $H_0 \sim 500 \text{ kms}^{-1} \text{Mpc}^{-1} : D^2 \sim 10^{-55} \text{ cm}^{-2}$

★ Age estimate problematic

Age from Friedmann (10^{10} yr)

■ Not a periodic solution

"Model fails at P = 0"



Oxford lecture (May 1931)

$$D = \frac{1}{c} \frac{1}{l} \frac{dl}{dt} = \frac{1}{c} \frac{1}{P} \frac{dP}{dt}$$

$$D^{2} = \frac{1}{P^{2}} \frac{P_{0} - P}{P} \sim \frac{1}{P^{2}} \qquad (1a)$$

$$D^{2} \times \frac{R_{0}}{S} \frac{P_{0} - P}{P} \sim \frac{1}{R_{0}} \qquad (2a)$$

$$D^{2} \sim 10^{-53}$$

$$C \sim 10^{-26}$$

$$P \sim 10^{8} \text{ GiV} \qquad (10^{11}) \text{ T}$$

$$t \sim 10^{10} (10^{11}) \text{ T}$$

Einstein-deSitter model (1932)

Remove curvature (Occam's razor)

All curvatures possible (Heckmann)

♯ Adopt Friedmann analysis

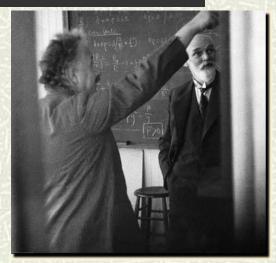
Time-varying universe with k = 0, $\lambda = 0$ Critical universe

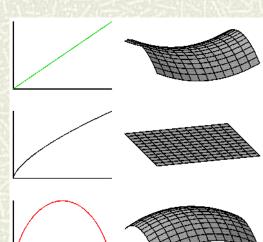
Calculate critical density

10-28 g/cm³: agrees with astrophysics

★ Well-known model

Despite age problem





An origin for the universe? (1931)

Rewind Hubble graph

U smaller in the past

Extremely dense, extremely hot

Primeval atom
Expanding and cooling since

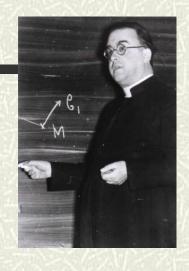
Singularity problem

 ∞ density, ∞ temp at t = 0?

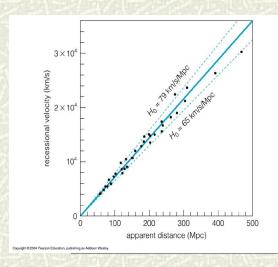
Quantum theory

Age problem

U younger than stars?



The big bang



Lemaître's hesitating universe (1931-34)

Primeval atom

Explosive expansion from radioactive decay

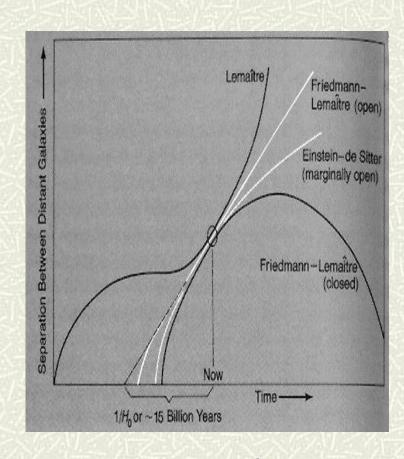
♯ Expansion slows down

Positive cosmic constant Energy of vacuum; stagnation

♯ Indefinite timespan

No age problem Formation of structure?

★ Accelerated expansionde Sitter universe at large t



Cosmic rays = radiation from early universe?

Formation of galaxies?

Growth in static universe

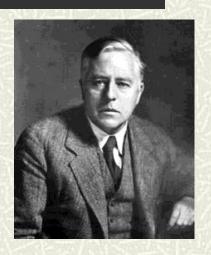
Natural fluctuations in density Exponential growth by gravitational collapse $\lambda_i = c_s/(G\rho_0/\pi)^{1/2}$

Growth in expanding universe

Lemaître 1934, Tolman 1935 Linear growth of density perturbations $\delta \rho / \rho \propto R$

Structure not from density fluctuations?

New mechanism needed Eddington-Lemaître model?



James Jeans





Observational parameters needed

$$k = -1, 0, 1$$
?

$$\lambda = 0$$
?

$$q_0 = -\ddot{R}/\dot{R}^2$$

■ Density of matter

$$\rho < \rho_{crit}$$
 ?

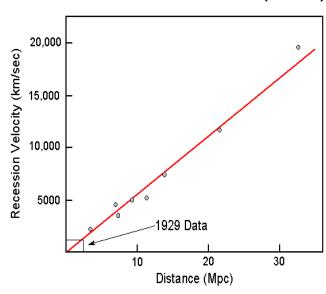
Hubble constant

$$\dot{R}/R = 500 \, km s^{-1} Mpc^{-1}$$
?

Timespan

$$\tau = 10^9 \text{ yr ?}$$

Hubble & Humason (1931)



What do redshifts represent? Is expansion a local effect?

Hubble and Tolman 1935

Paradigm shift or slow dawning?

Hubble/Slipher Empirical law for nebulae

Friedmann Time-varying solutions

Lemaître Theory and observation

Models de Sitter, Eddington, Einstein,

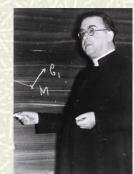
Tolman, Robertson

Slow emergence of theory and evidence











Bonus: Einstein's steady-state model

■ Non-static line element (1931)

$$ds^2 = -e^{2kt}(dx_1^2 + dx_2^2 + dx_3^2) + c^2dt^2$$

♯ Age problem

Conflict with stellar ages

♯ Non-evolving universe

Continuous formation of matter Associated with λ : energy of space

Anticipates Hoyle et al.

Fundamental flaw
Not published

Le Titter much Tolman haben breits gezeigt, class es Tisenigen der Gleichungen (1) gibt, welche den Beobachtungen gerecht werden. Es erzub sich aber die Teluvierigkeit, dass die Theorie stets auf einen zeitlschen Aufang fithrite, won der Gebensen Gedenung 10 10 - 10 " Jahre zurückliegt, was aus verschiedenen Ynsuden unannehmber sehden.

In Nachfolgenden will sich auf eine Lösung der Gleichung (1) aufmerksam machen, melche Hubbel's Thatsuchen gerecht wird, und in welcher die Bielete zeitlich konstant est. Ihre-Lösung ist zwar in dem ullzuminen Schema Tolman's unthalten, sehnant über hisher wicht in Betracht gezogen worden zu sein. 1 Ich setze au

Einstein's exploration of a steady-state model

♯ Why does model fail?

De Sitter model (9/4 \rightarrow -3/4)

$$\rho = 0$$

How is matter formed?

No 'creation' term

Einstein's crossroads

Realised S-S model requires term
Declined to add term to GFE

Evolving models

Less contrived

Set
$$\lambda = 0$$

Die Gleichungen (1) leeform $-\frac{3}{4}\alpha^2 + dc^2 = 0$ $\frac{3}{4}\alpha^2 - dc^2 = xp c^2$

reduce
$$\alpha^2 = \frac{1}{3} \varphi \epsilon^2 \frac{\kappa c^2}{3} \varphi \cdots (4)$$

Die Bielete ist also konstant und bestimmt die Tapansion Les auf das Vorgeichen.

Retrustret man ein durch physische Mussetisbe belgrengtes Valumen, so wandert managesetzt materielle Teilchen aus demselben heners, Danist die Fielite konstant blibe, misse immer neue Mussenteilchen in dem Blumen aus dem Ramme entstehen.

Der Erhaltungssatz bleebt deedurch zuwahrt, dass bei Setzung des 2-Gleedes dur Ramm selbst micht energetisch leer est; seine Geltung wird bekanntlisch durch des Gleechungens(1) gewährleistet. Jun Nachfolgenden will ich auf eine Lösseng der Gleichung

(1) aufmerkesem machen, welche Hubbel's Thatsuchen gerecht
wird, und in welcher die Tielete zeitlich konstant est. ItereLösung ist zwar in dem allzumeinen Schema Tolman's urthalten,
scheimt über hisher wicht in Betracht zezogen worden zu seen.

1 Ich setze au

ds2 =- e (dx,2+ dx,2+ dx,2) + c2dt2 ... (2)

Die Bielite ist also konstant und bertimmet die Tapansion bes auf des Vorgeichen.

Retrachtet man ein durch physische Musstribe begreugtes Valumen, so wandert managesetzt materielle Teilchen aus demselben henans, Danist des Fiehte konstant bleibe, misse immer neue Massenteilchen in dem Volumen aus dem Ramme entstehen.

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Coda: Slow acceptance (1935-65)

Hot big bang (1940s)

Nucleosynthesis in the hot infant universe? Background radiation from early universe?

Little interest from community

No search for the cosmic radiation General relativity difficult, abstruse

■ Steady-state universe (1948)

Continuous creation of matter from vacuum No age or singularity problems

Later ruled out by experiment (1960s) *Radio-galaxy counts, CMB*



Gamow, Alpher and Hermann

Hoyle, Bondi and Gold



Paradigm shift or slow dawning?

Revolutionary v normal science

Normal science interspersed by revolutions

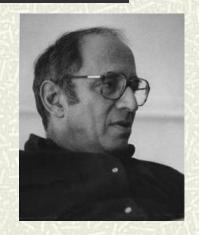
■ The paradigm shift

Change of worldview occurs
Social factors important

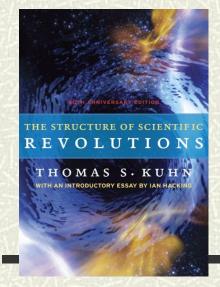
♯ Incommensurability

New worldview incommensurate with old

Exp U: Slow exploration of theory and observation Slow acceptance of new paradigm (1935-65)



Thomas Kuhn



The big bang model - questions

Nature of dark energy?

Role in BB?

***** Nature of dark matter?

Particle experiments?

Which model of inflation?

The multiverse?

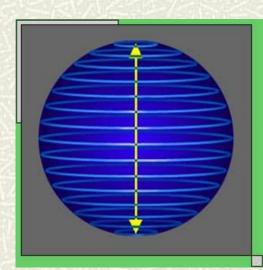
♯ The singularity problem

What banged?

What does time zero mean?



No-boundary universe



The case is never closed

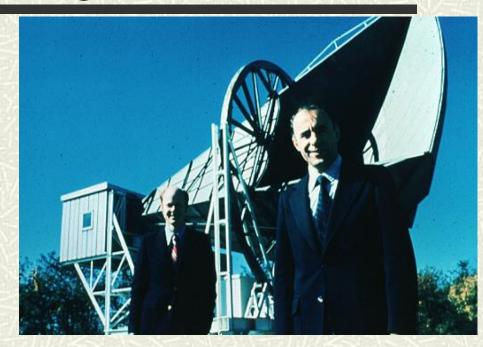
Act V Cosmic background radiation

- **♯** Search for radio signals

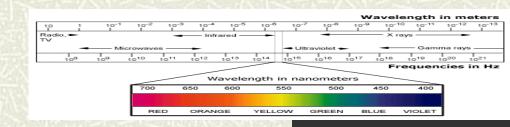
 Large, sensitive receiver
- □ Universal signal (1965)
 From every direction
- **Low frequency (microwave)** *Low temperature (3K)*
- # Echo of big bang

 Radiation from early universe

BB model goes mainstream



Penzias and Wilson



Cosmology today

Satellite measurements of CMB

No interference from atmosphere

Expected temperature

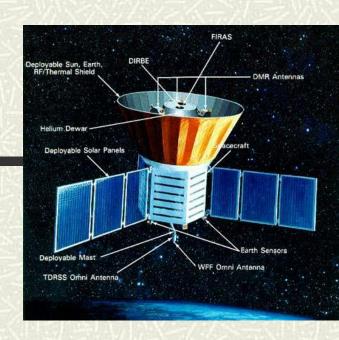
Expected frequency

Full spectrum

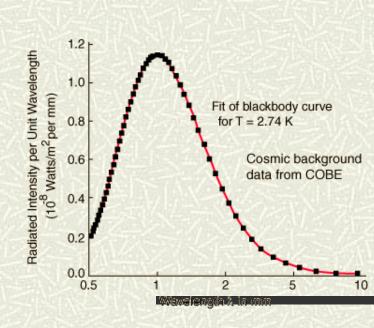
Perfect blackbody spectrum

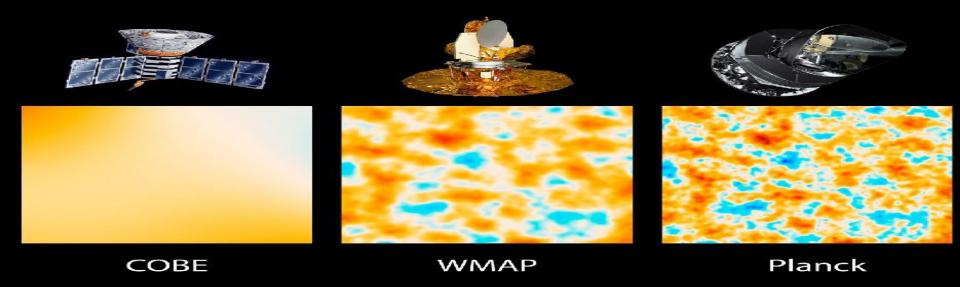
Perturbations

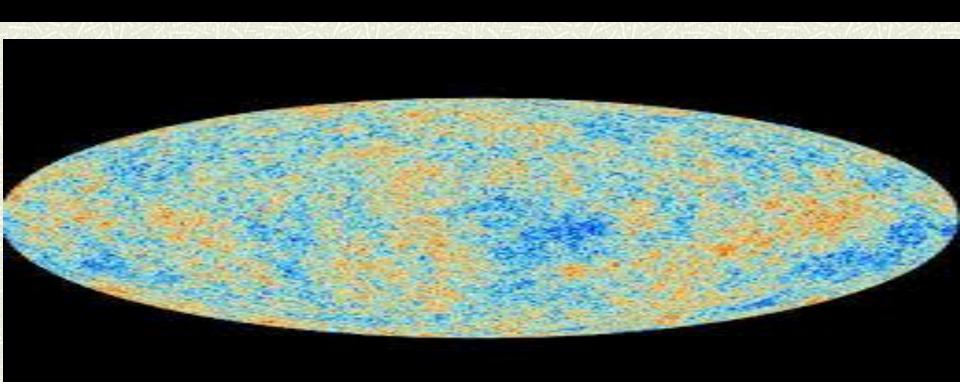
Variation of 1 in 10⁵



COBE satellite (1992)







New results: Planck Satellite (ESA, 2013)

1. Improved sensitivity

 $\Delta T/T \approx 1 \times 10^{-6}$

2. Full spectrum of *T* anisotropy

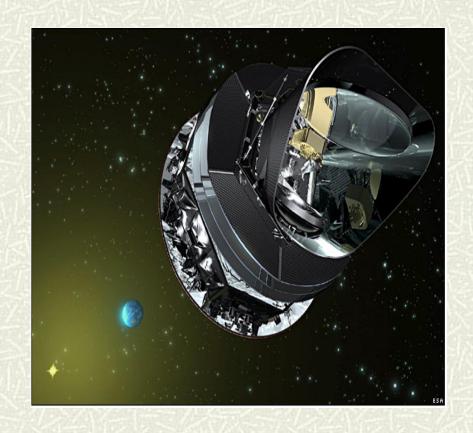
New acoustic peaks :scale invariance? Accurate values for Ω_{Λ} , Ω_{M}

3. Gravitational lensing

Remove degeneracies

4. Polarization measurements

E-modes: fluctuations B-modes: gravity waves?



Planck results (2013)

1. New Hubble constant

67.3 + -1.2 km/s/MPC Age = 13.8 billion yr

No age conflict with astrophysics

2. Curvature: flat

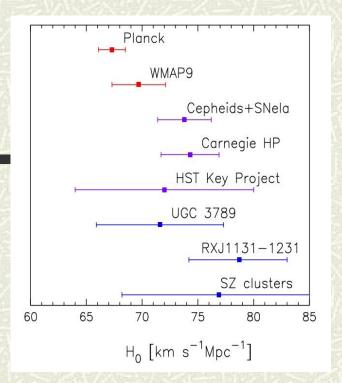
$$\Omega_{\rm k} = -0.0005 + -.07$$

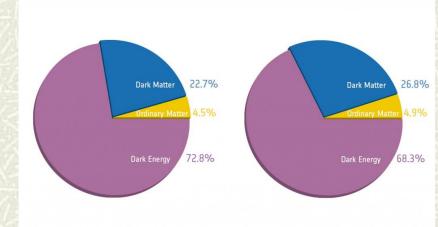
3. Positive cosmic constant

$$\Omega_{\Lambda} = 68\%$$

4. New mass/energy parameters

$$\Omega_{\rm DM} = 27\%$$
, $\Omega_{\rm OM} = 4.9 \%$





After Planck

Before Planck

Planck Results

1. Power spectrum

Not scale invariant $n_s = 0.96$

2. Compatible with inflation

Simple 'slow-roll' models Higgs-type field?

3. Complex inflation out

Double field out Hybrid models out Cyclic models out

