# New perspectives on steady-state cosmology

From Einstein to Hoyle

Cormac O'Raifeartaigh (Waterford Inst of Technology)
Simon Mitton (University of Cambridge)

## Einstein's steady-state model

#### **♯** Unpublished AE manuscript

Written in early 1931

#### **♯** Contains 'steady-state' model of the cosmos

Expanding universe of constant matter density Continuous formation of matter from vacuum Anticipates controversial theory (Hoyle)

#### **#** Inconsistent model

Fatal flaw Abandoned, not amended

#### **#** Evolving models embraced

Friedman-Einstein, Einstein-de Sitter



Einstein in California (1931)

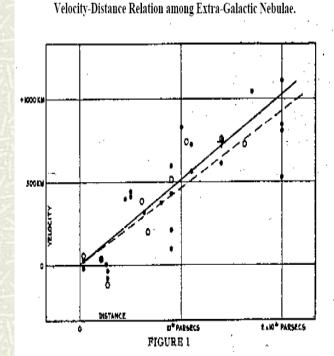






- **♯** Spiral nebulae extra-galactic (1925)
  - 100-inch reflector at Mt Wilson Resolved Cepheid stars in several nebulae Leavitt's period-luminosity relation
- **A distance/redshift relation for the spirals?**Redshifts of the nebulae by VM Slipher (1915,1917)
- **Approx linear relation (Hubble, 1929)**

Some anomalies (Peacock)
Slipher not acknowledged



$$H = 585 \text{ kms}^{-1} \text{Mpc}^{-1}$$

## A cosmic puzzle

**#** What is causing recession of the galaxies?

If redshifts are velocities
If effect is non-local

**■** Newton's law of gravity

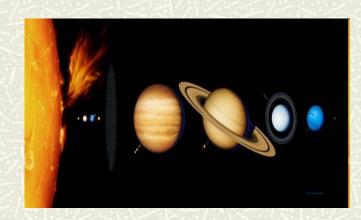
Gravity pulls in, not out

No other long range force for neutral matter

**♯** Space, time are fixed

Not affected by contents of universe Eternal, infinite universe





## General relativity (1915)

**♯** Space+ time = space-time

Space-time dynamic

Distorted by motion, mass

**Gravity** = curvature of space-time

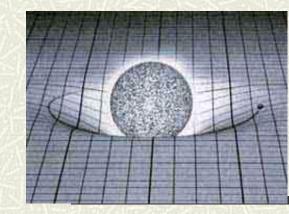
$$G_{\mu\nu} = \frac{8\pi G}{c^4} T_{\mu\nu}$$

**#** Empirical evidence

Perihelion of Mercury
Bending of starlight (Eddington, 1919)



Albert Einstein

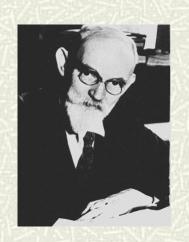


## Einstein's universe (1917)

- **#** Apply general relativity to the cosmos
- **Expanding or contracting**
- **No evidence for such a universe**Unaware of Slipher redshifts
- **Add cosmic constant to give 'static' solution**

$$G_{\mu\nu} + \lambda g_{\mu\nu} = \frac{8\pi G}{c^4} T_{\mu\nu}$$





## Friedman models of the cosmos

**★ Allow time-varying solutions to the field equations**Expanding, contracting universes



**#** Geometry, evolution depends on matter content

Positive curvature (1922) Hyperbolic curvature (1924)

**Evolving models (Zf. Ph.)** 

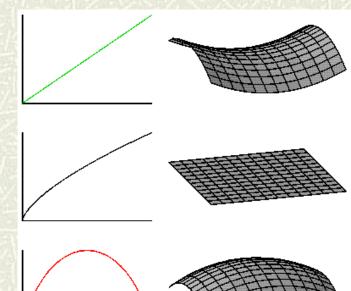
Matter density varies over time

**■ Ignored by community** 

Disliked by Einstein

Correction and retraction

Alexander Friedman 1888 -1925



## Lemaître's universe (1927)



Rate of expansion from mean distances and redshifts  $H = 585 \text{ km/s/Mpc} \quad (1927)$ 

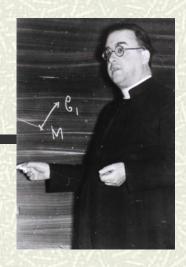


Starts from Einstein universe at  $t = -\infty$ 

#### **♯** Rejected by Einstein

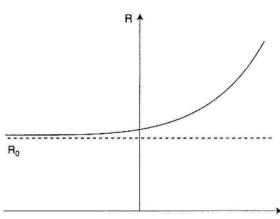
"Votre physique est abominable"

Ditto for Friedman models



Fr Georges Lemaître

Not an empirical law



## An expanding universe? (1930-)

#### **RAS** meeting (1930)

Eddington, de Sitter

<u>If</u> redshifts are velocities, and effect is non-local

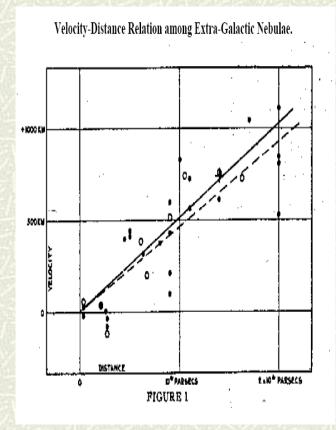
# • Hubble's law = expansion of space? Static relativistic models don't fit data

Static relativistic models don't fit data
Dynamic models required

#### Friedman-Lemaître models

Time-varying radius
Variable matter density

<u>Evolving universe</u>



Cosmic expansion?

## The expanding, evolving universe (1930 -)

- **Eddington** (1930, 31)
  - On the instability of the Einstein universe Expansion caused by condensation?
- de Sitter (1930, 31)

Further remarks on the expanding universe Expanding universes of every flavour

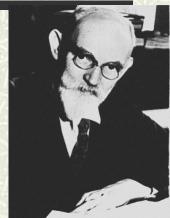
■ Tolman (1930, 31)

On the behaviour of non-static models Expansion caused by annihilation of matter?

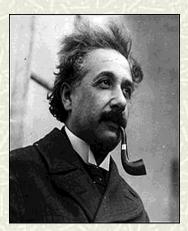
• Einstein (1931, 32)

Friedman-Einstein model  $\lambda = 0, k = 1$ Einstein-de Sitter model  $\lambda = 0, k = 0$ 









If redshifts represent expansion...

Evolving models

# New: Einstein's steady-state model (1931?)

© The Hebrew University of Jerusalem

- **Filed** as draft of *F-E* model Similar title, opening
- **#** Cites Hubble's law Cites instability of static model
- **#** Cites evolving models Discusses age problem
- **Proposes alternative solution** Expanding, unchanging cosmos? Continuous creation of matter Associates with  $\lambda$  - energy of space

Zum kosmologischen Troblem. A. Binstein . The wichtigste grundsestzliche Schwierigkeit, welche sich zeigt, nem man nach der Art fragt, we die Materse der Gen Raum in sehr grossen Dimensionen afsellt, liegt

bekanntlich durin, duss die Granitationsgesetze im Algemeiner

met der Hypothese einer endlichen milleren Tichte der Katerie wicht verträglich sind. Schon zu der Zeit, als man woch allymen an Newtons Gravitations - Theorie fisthielt, hat deshalt Teeliger das Newton sehr Gesetz durch tong tostands-

tunktion modefiziert, welche für grosse Hostiende ve erheblich

האוניברסיטה העברית בירושלים

Die Gleichungen (1) Vieform - 1 a2 + de2 = 0 3 x2 - 1c2=xpc2 rder Die Gielete ist also konstant und bestimmt die Tapansion

schneller abfallt als 72.

bes and dus Porgeichen.

## Einstein's steady-state model: key quotes

#### **New solution**

"In what follows, I wish to draw attention to a solution to equation (1) that can account for Hubbel's facts, and in which the density is constant over time"

#### **Matter creation**

"If one considers a physically bounded volume, particles of matter will be continually leaving it. For the density to remain constant, new particles of matter must be continually formed within that volume from space "

#### Dark energy

"The conservation law is preserved in that, by setting the  $\lambda$ -term, space itself is not empty of energy; its validity is well known to be guaranteed by equations (1)."

## An abandoned model

# A fatal flaw

De Sitter metric

Matter creation associated with λ

- **Null result masked by error**Derivation incorrect
- **Einstein's crossroads**Identifed problem on revisionDeclined to amend GFE
- **#** Evolving models?

*Less contrived and set*  $\lambda = 0$ 

Jun Nachfolgenden will ich auf eine Lösung der Gleichung (4) aufmerksehm machen, welche Hubbel's Thatsuchen gerecht wird, und in welcher die Bielete zeitlich konstant est. Dere Lisung ist zwar in dem allzemeinen Schema Tohnan's urthalten, sehnant aber hisher wielet in Betracht zezogen worden zu seen.

Die Gleichungen (1) before  $-\frac{3}{4}x^2 + dc^2 = 0$   $\frac{3}{4}x^2 - dc^2 = xpc^2$ 

reduce 
$$\alpha^2 = \frac{1}{3} \varphi e^2 \frac{\kappa c^2}{3} \varphi \dots (4)$$

Die Dielite ist also konstant und bestimmt die Zepansion Les auf das Porgeichen.

## The steady-state universe (1948)

#### **Expanding but unchanging universe**

Hoyle, Bondi and Gold (1948)

No beginning, no age paradox

No assumptions about physics of early epochs

#### **#** Continuous creation of matter

Very little matter required

#### **\blacksquare** Replace $\lambda$ with creation term (Hoyle)

 $G_{\mu\nu} + C_{\mu\nu} = k T_{\mu\nu}$ Conservation of energy violated

#### **■ Improved version (1962)**





Bondi, Gold and Hoyle



Hoyle and Narlikar (1962)

## Evolving vs steady-state universe

#### **#** Radio-astronomy

Galaxy distributions at different epochs
Cambridge 3C Survey (Ryle)

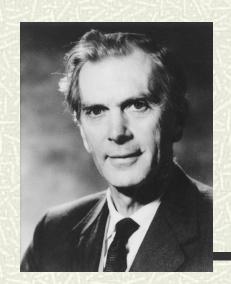
#### **#** Cosmic microwave background

Low temperature, low frequency Remnant of early universe

#### **#** Optical astronomy

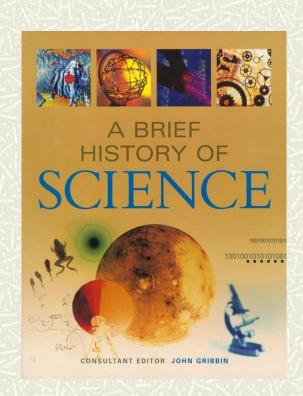
Amended timescale of expansion (Baade, Sandage)





## Significance of Einstein's steady-state model

- □ Unsuccessful theories important
   Understanding the development of successful theories
- **New perspective on steady-state theory**Logical possibility: not a crank theory
- Insight into Einstein's philosophy
   Discards model rather than add new term to GFE
   Occam's razor approach
- Insight into scientific progress
   Not Kuhnian paradigm shift
   Slow dawning



Links with modern cosmology Dark energy: creation energy and  $\lambda$ Cosmic inflation: de Sitter metric



NATURE | NEWS

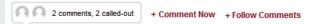
### Einstein's lost theory uncovered

Physicist explored the idea of a steady-state Universe in 1931.

Davide Castelvecchi

24 February 2014

# New Discovery Reveals Einsteir Tried To Devise A Steady State Model Of The Universe



Almost 20 years before the late Fred Hoyle and his colleagues devised the <u>Steady State Theory</u>, Albert Einstein toyed with a similar idea: that the universe was eternal, expanding outward with a consistent input of spontaneously generating matter.

An Irish physicist came across the paper last year and could hardly believe According to this week's article in <u>Nature</u>,

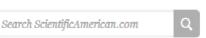
model of the universe very different to today's Big Bang Theory.

The manuscript, which hadn't been referred to by scientists for decades.





## SCIENTIFIC AMERICAN™

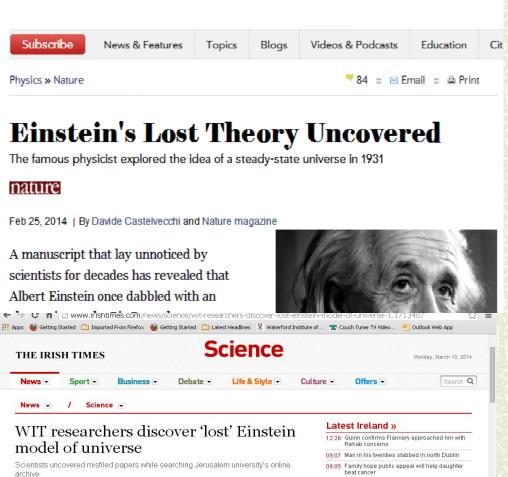


08:42 Gardaí investigate death of woman in Dublin

08:25 Flannery faces call from all parties to attend

The way back isn't so simple

Sign In | Register 🔲 0



## Einstein's steady-state model and cosmology today

#### **#** Dark energy (1998)

Accelerated expansion (observation)
Positive cosmological constant

#### **♯** Einstein's dark energy

"The conservation law is preserved in that, by setting the  $\lambda$ -term, space itself is not empty of energy; its validity is well known to be guaranteed by equations (1)."

#### **#** Cosmic inflation

Inflationary models use de Sitter metric
Used in all steady-state models
Flat curvature, constant rate of matter creation
Different time-frame!

